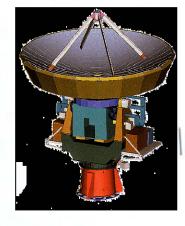
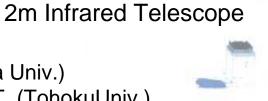
Current status of Japanese activities for astronomy in Antarctica

Takashi Ichikawa & Japanese Consortium



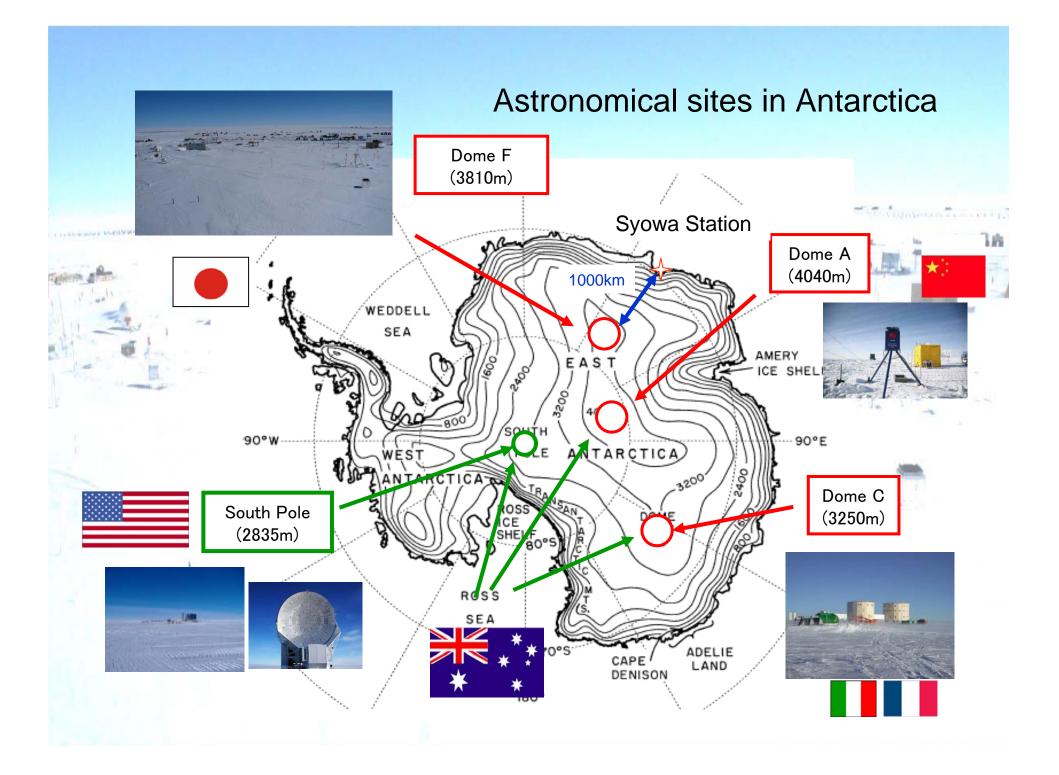


(Chair) Nakai, N., Seta M. (Tsukuba Univ.) Ichikawa, T., Okano, S., Sakamoi, T. (TohokuUniv.) Taguchi, M. (Rikyou Univ.) Takato, N., Uraguchi, H., Iye, M. (NOAJ) Kurita, M. (Nagoya Univ.) Motoyama, H. (NIPRJ)

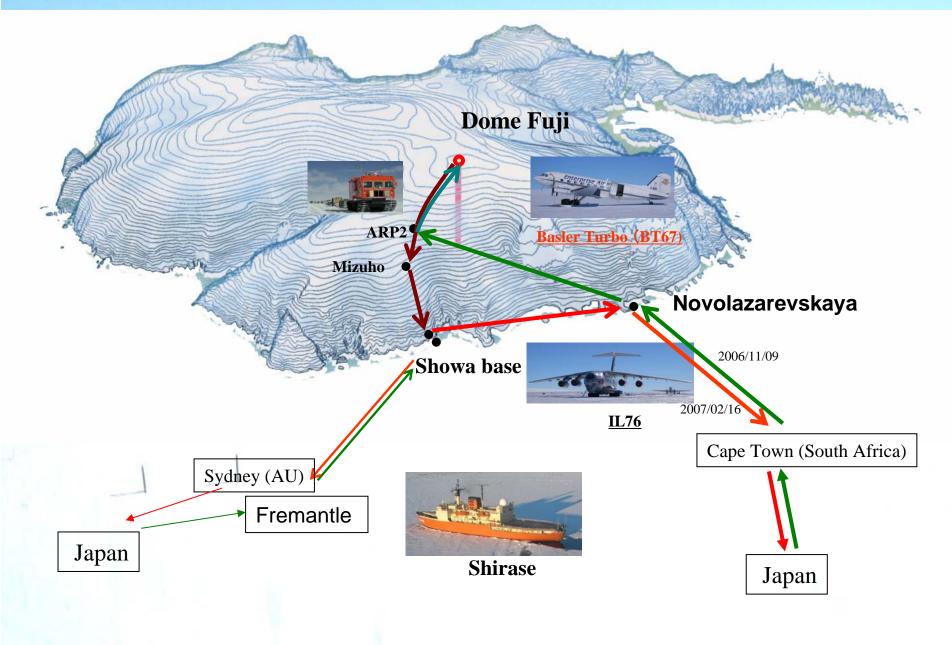


THz Radio Telescope

& collaborators



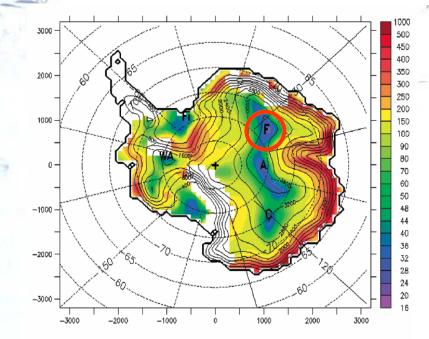
A long way to Dome Fuji



Japan has one of best astronomical sites on earth

Clear sky (photometric day > ~85%) Low and stable humidity (PWV < 0.6mm) Low temperature (-70°C in winter) Good seeing above boundary layer weak wind

According to model atmosphere, the surface boundary layer at Dome Fuji is expected to be thinnest in Antarctic inland



Height where the residual boundary layer seeing is 0.1'' or better 50% of the time (JJA 2004)

Swain & Gallee (2006)

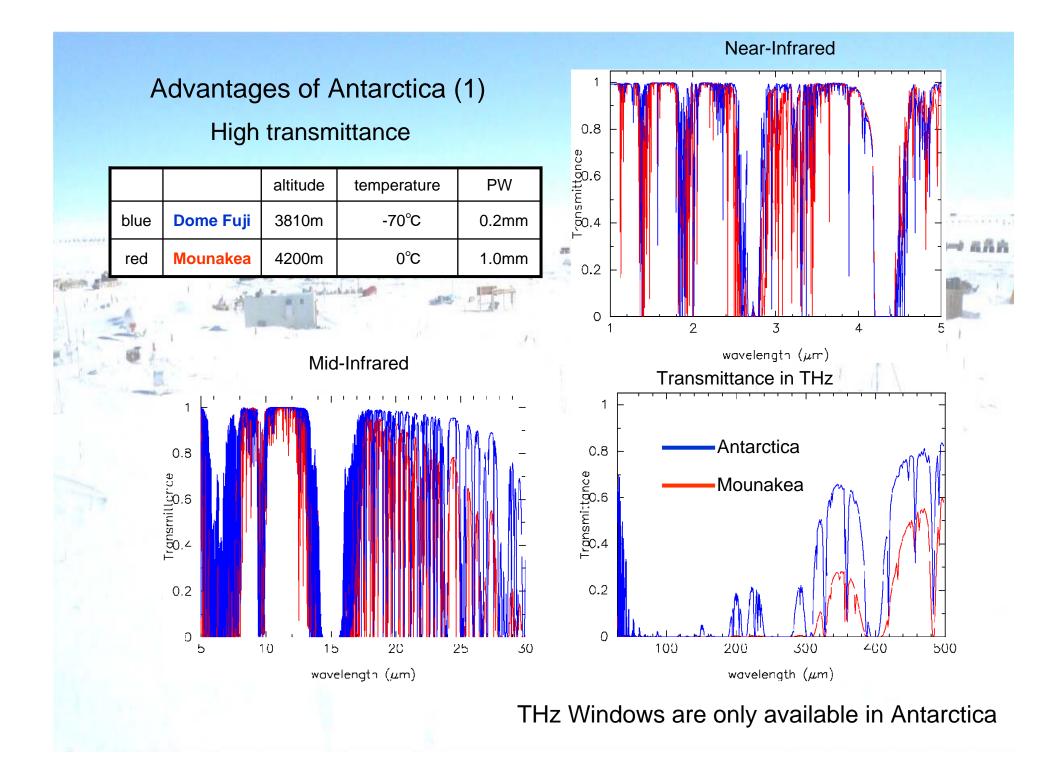
The seeing above boundary layer (~30m) is ~0.3"-0.4" (λ =0.5µm) in winter at Dome C

Agabi+ (2005) Lawrence+ (2006)



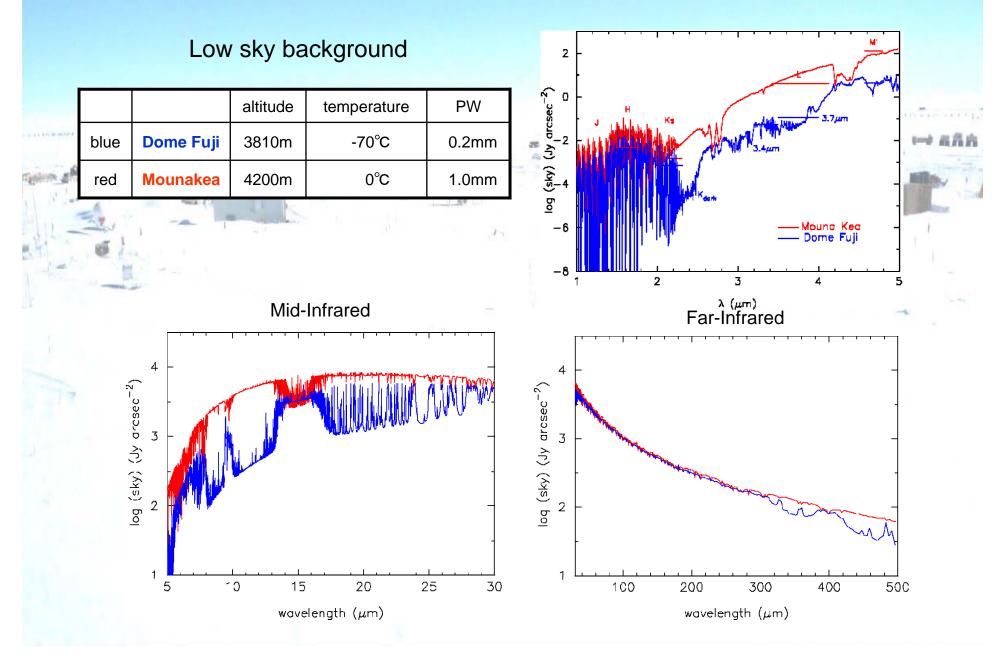
At Dome Fuji, it is ~0.3"-0.4" at ~20m or above (?)

Better seeing in the daytime



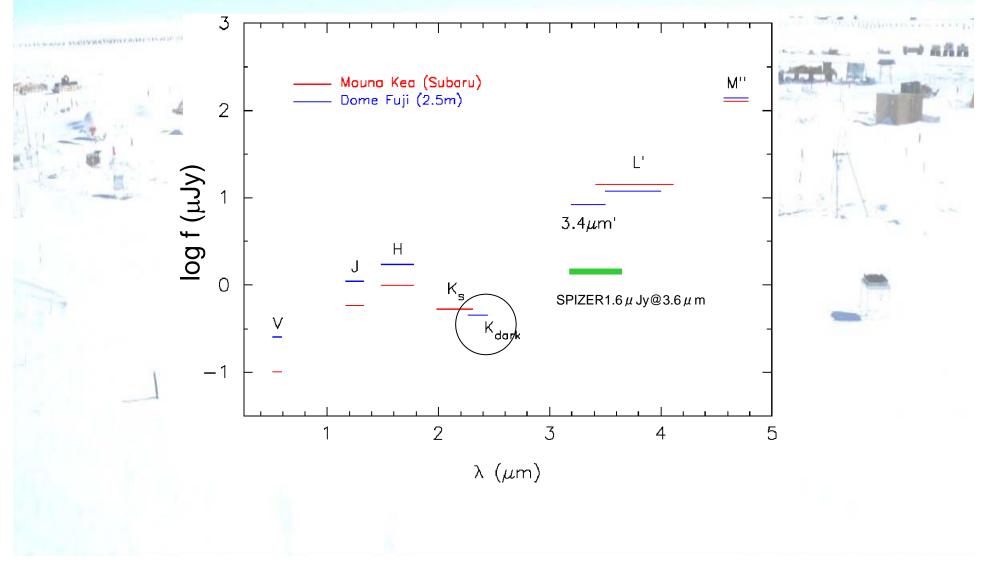
Advantages of Antarctica (2)

Near-Infrared

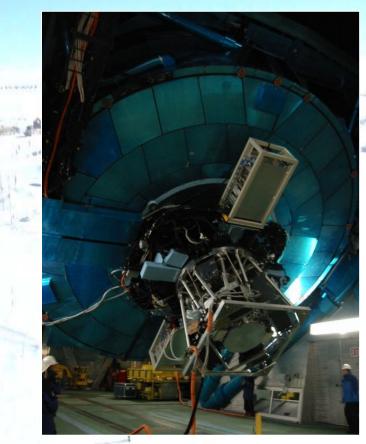


Comparison with Subaru for 2.5m Antarctic Telescope

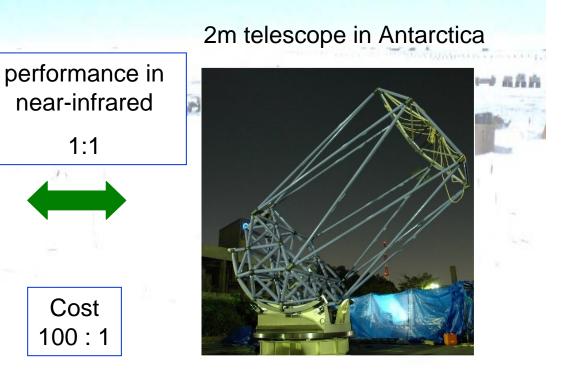
1 hour integration with S/N=5 σ for pint source



Subaru+MOIRCS



Ichikawa et al. (2005)



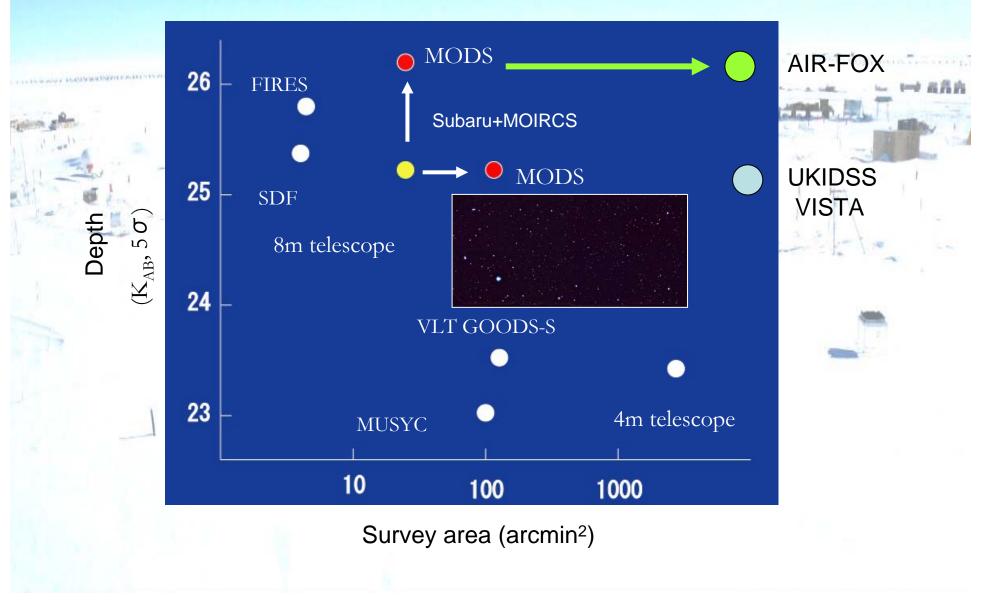
Kurita et al. (2005)

Ultra light weight mount

MOIRCS High Redshift Galaxy Survey In GOODS region ------22:00.0 28 hour integration 12:00.0 62:10:00.0 38:00.0 30.0 36:00.0 08:00.0 06:00.0

Widest and Deepest High-Redshift Galxy Survey

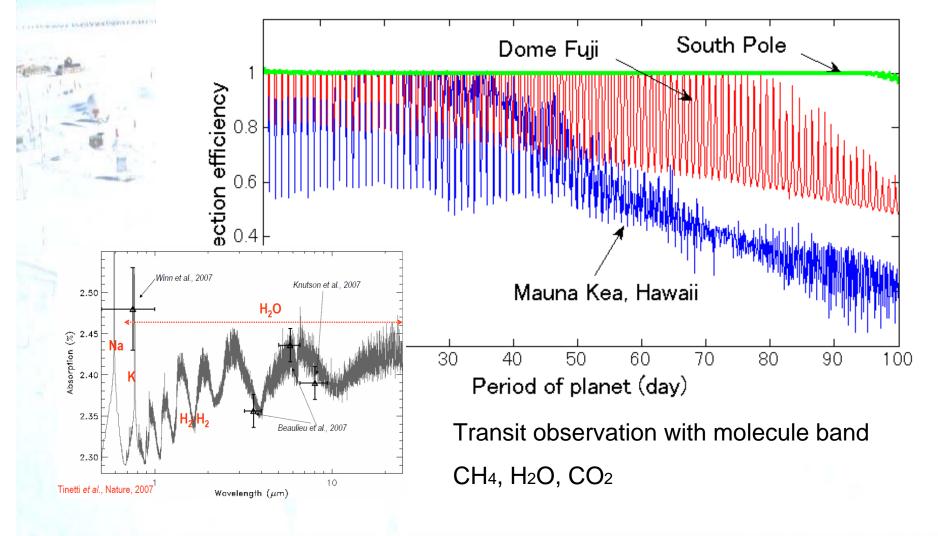
~2m telescope + (K_{dark} , 3.4µm) camera



Advantages of Antarctica (3). Polar night

Transit identification of long-period extra-solar planets

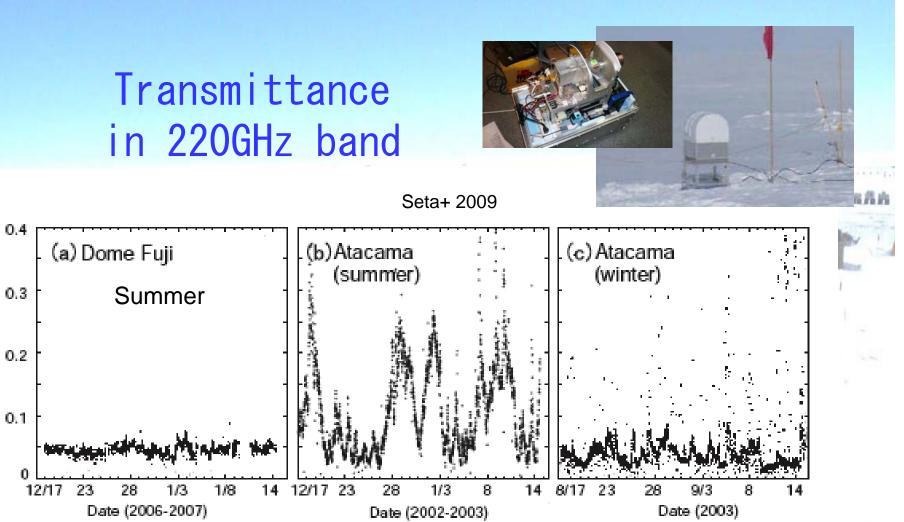
Probability of chance for multiple observation of transit (1-year monitoring)



Dome Fuji station

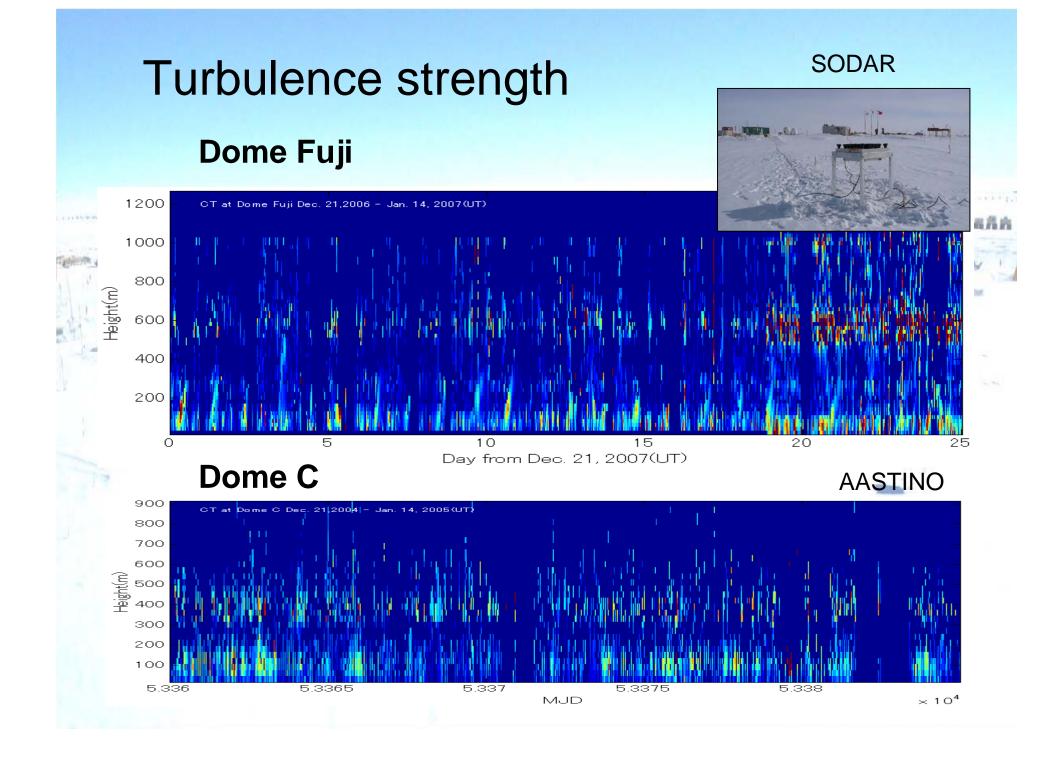
National Institute of Polar Research



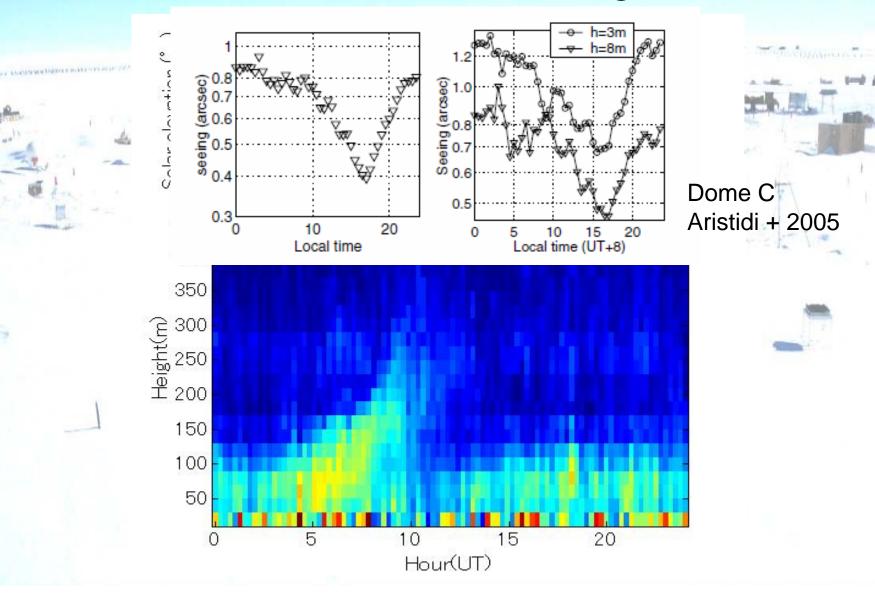


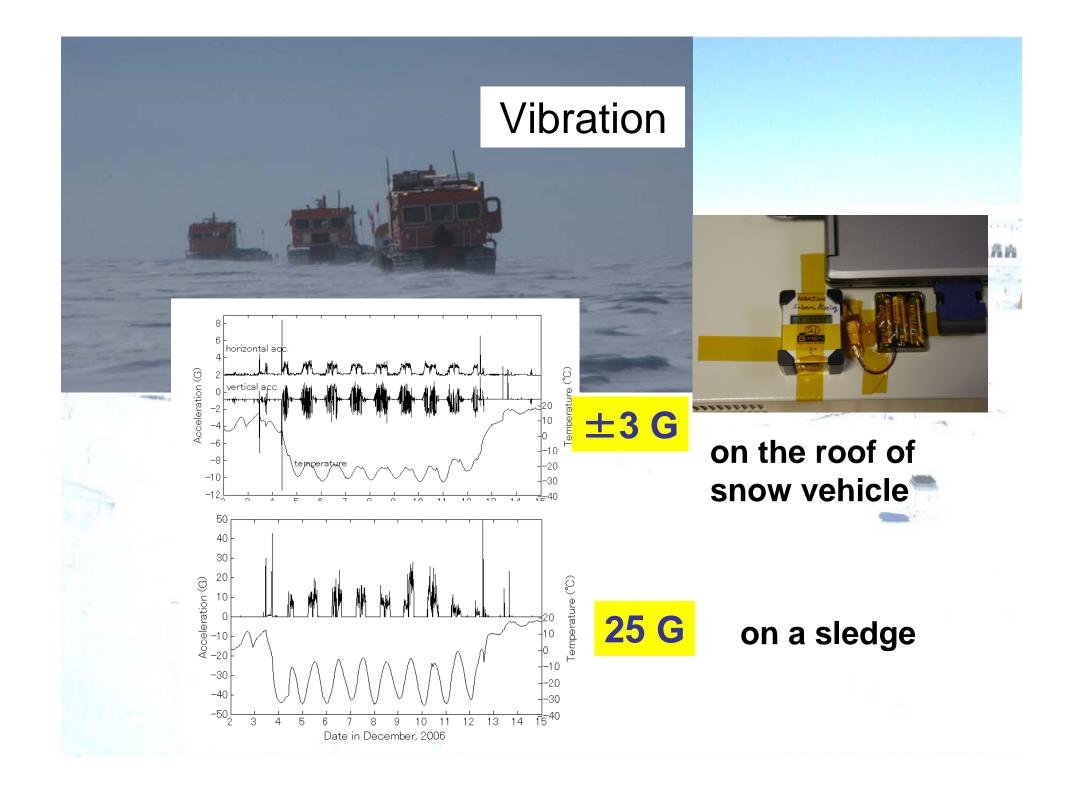
Very stable than Atakama in summer However, higher transmittance than Atakama in the best days

Optical Depth



Diurnal variation of turbulence strength





Pilot studies with small telescopes

J, K_{dark} Camera

40cm Infrared telescope

AIR-C

for seeing, transmittance, sky background

30cm THz telescope

for transmittance and its stability



Seta +



AIR-T-40 40 cm Antarctic Infra-Red Telescope

+ remote control (under developing)

Specs for the environment at -80°C

Operation at Hokkaido Rikybetsu the coldest place in Japan (2008/2) Murata et al.

deploy





assemble

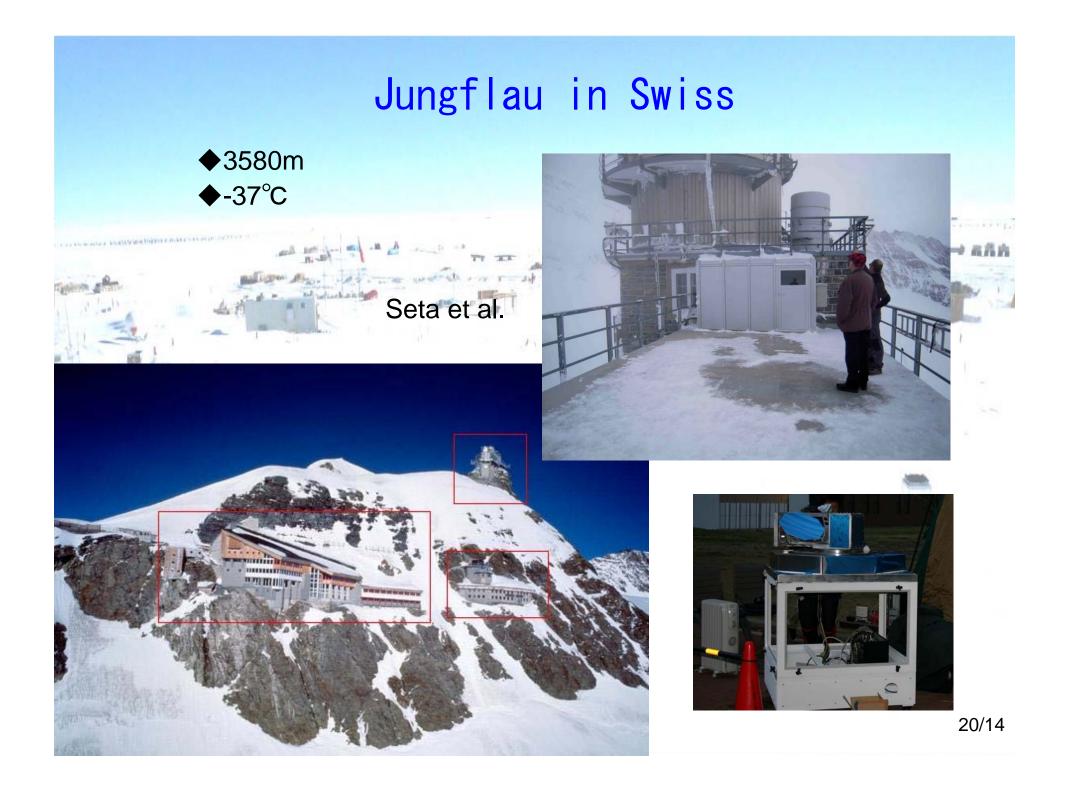


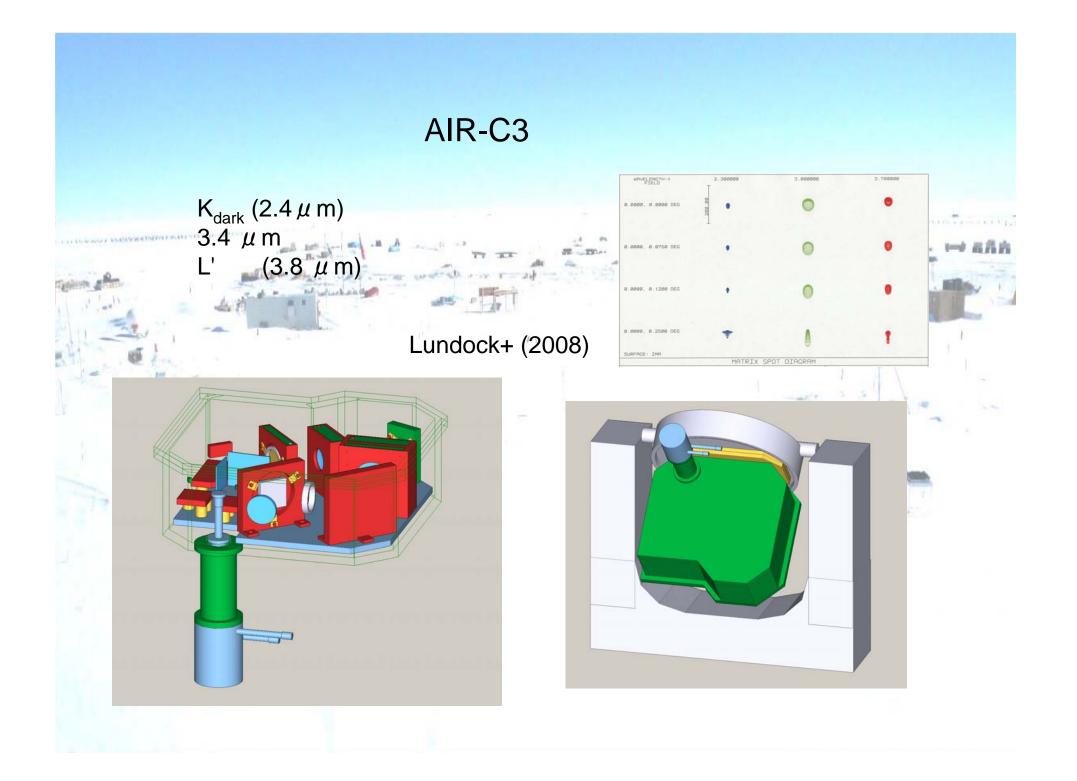


Operation at -23°C without no problems except note PC



Operation by wind power





Infrared Astronomy by Japanese group

Antarctic InfraRed astronomy at dome Fuji Observatory eXplorer

AIR-T-40 Antarctic InfraRed Telescope (40cm) AIR-C1 Antarctic InfraRed Camera 1 AIR-C3 Antarctic InfraRed Camera (3 bands) AIR-T-200 Antarctic InfraRed Telescope (2m)

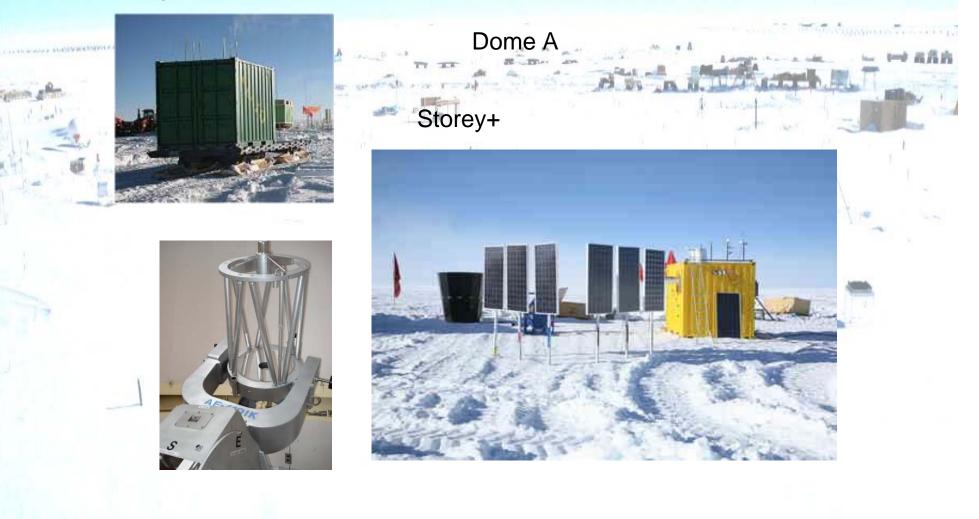
AIR-FO

AIR-Solar Antarctic InfraRed Solar telescope (40cm) AIR-HET10 Antarctic InfraRed Heterodyne Spectrograph (10 μ m) AIR-Trans Antarctic InfraRed Transit telescope etc.

 \leq 500 M yen (Telescope & Instruments) + NIPRJ (logistics)

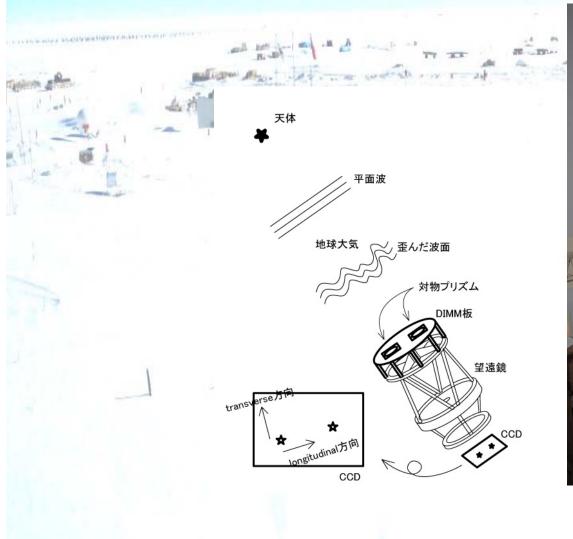
Collaboration with Australia group at Dome Fuji

Engine module



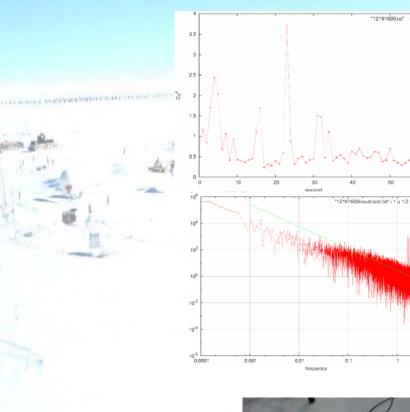
Seeing measurement by DIMM

Okita+





Seeing measurement by anemometer





Kurita et al.





Current status and future plane

2005/12 started the discussion with NIPR 2006/2 first observations for astronomical test

SODAR

220GHz radio meter

2009/12-2010/2

first step on Dome Fuji by a Japanese astronomer transmittance by handy infrared spectrograph transmittance by 200GHz radiometer

2010/12-2011/2

deployment of 40cm-infrared and 30cm-THz telescopes at Dome Fuji

2010-2014 Construction of overnight facilities by NIPR

- 2011-2012 deployment of PLATO-Fuji by collaboration with UNSW Observations with small telescope over winter (remote operation)
- 2014?- Construction of large telescope(?)

