

Dome-F, Antarctica

Antarctic Infrared Telescope (AIR-T) and its Scientific Drivers

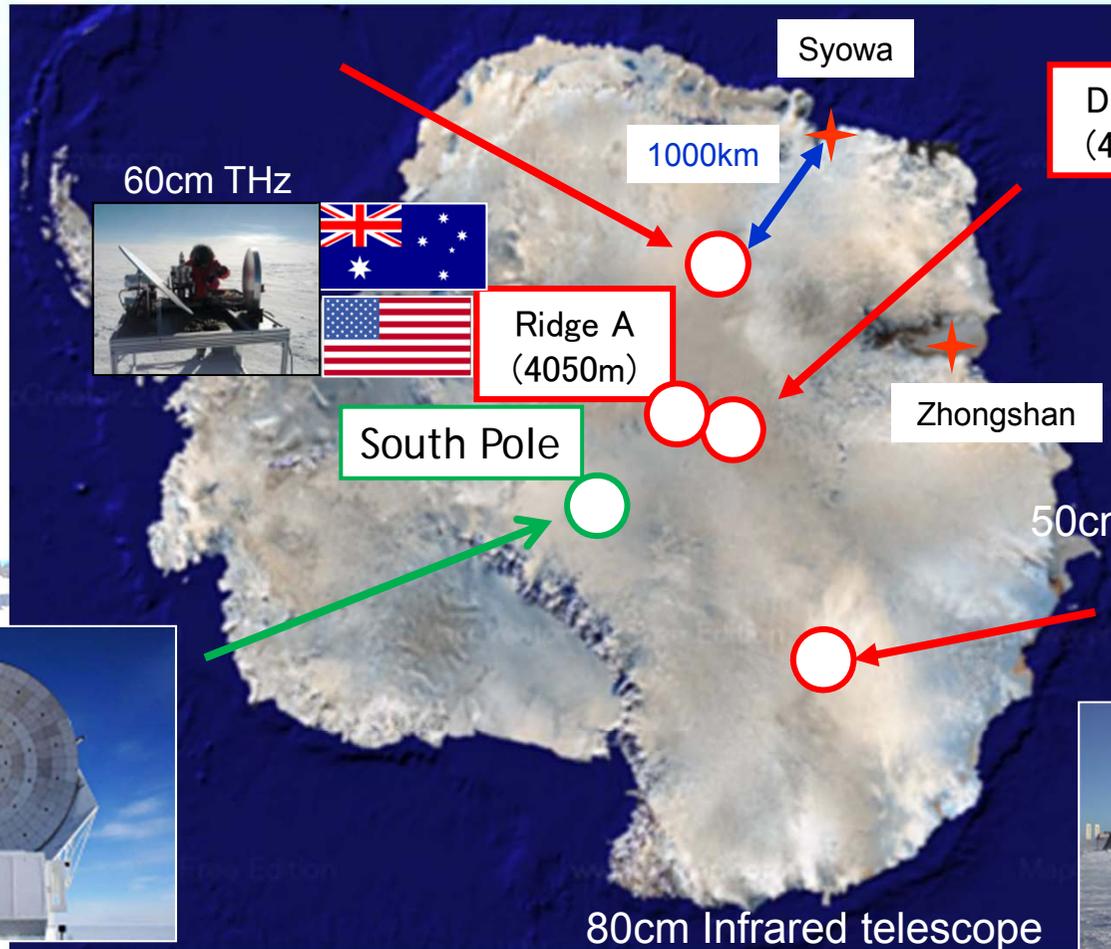
Tohoku University

Takashi Ichikawa

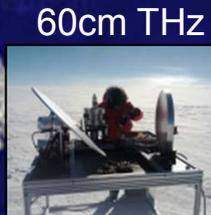
The ideal site for astronomy



Dome F
(3810m)



Dome A
(4100m)



60cm THz



Ridge A
(4050m)

South Pole

Zhongshan



50cm Schmidt telescope

Dome C
(3250m)



10m Sub-mm telescope

80cm Infrared telescope
40cm optical telescope



Astronomical Station at Dome-Fuji since 2011

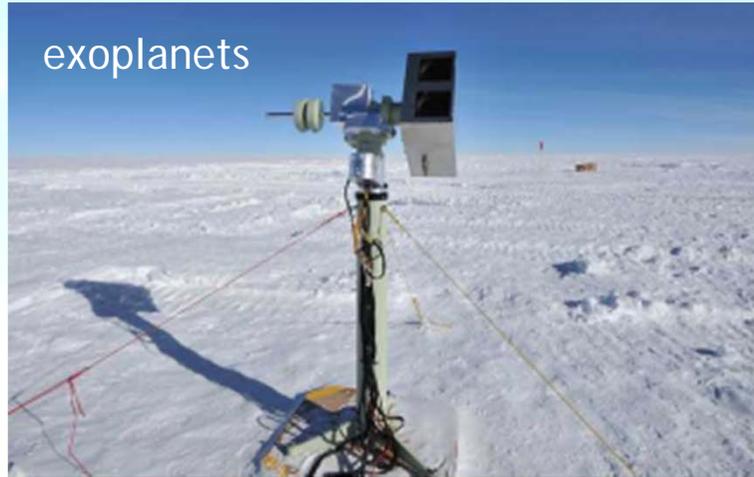
National Institute of Polar Research

Jan, 2013

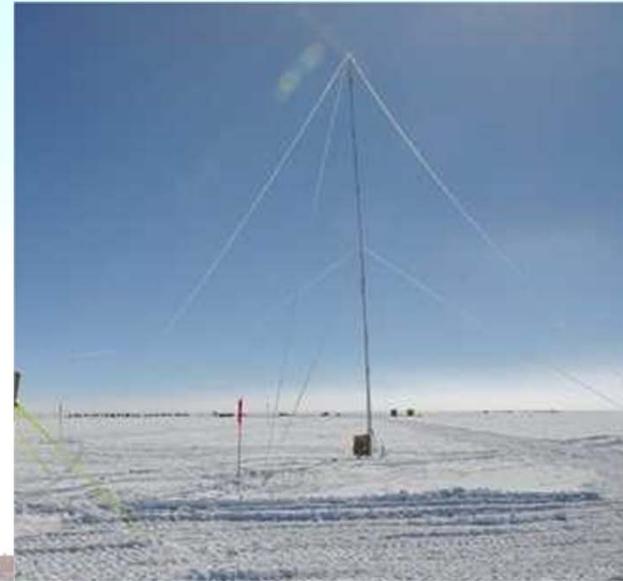


PLATO-F (UNSW)
Power station and instruments

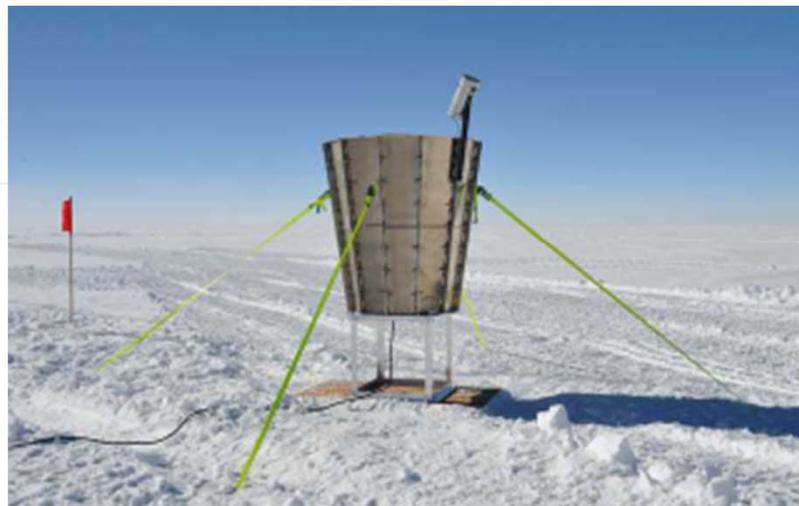
TwinCAM
10cmx2 telescopes



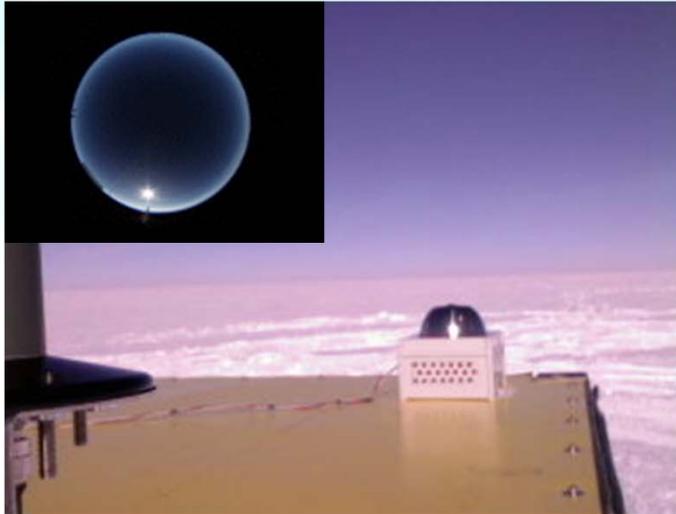
16m pole
temperature



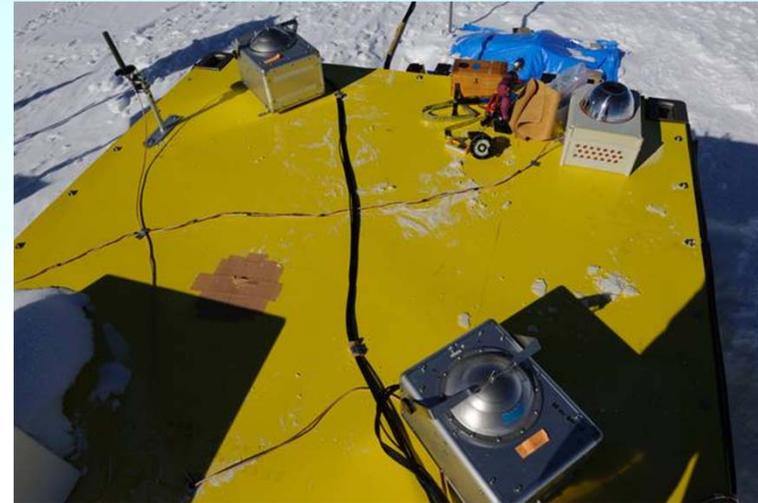
SNODAR (Bonner et al.)



Whole sky camera

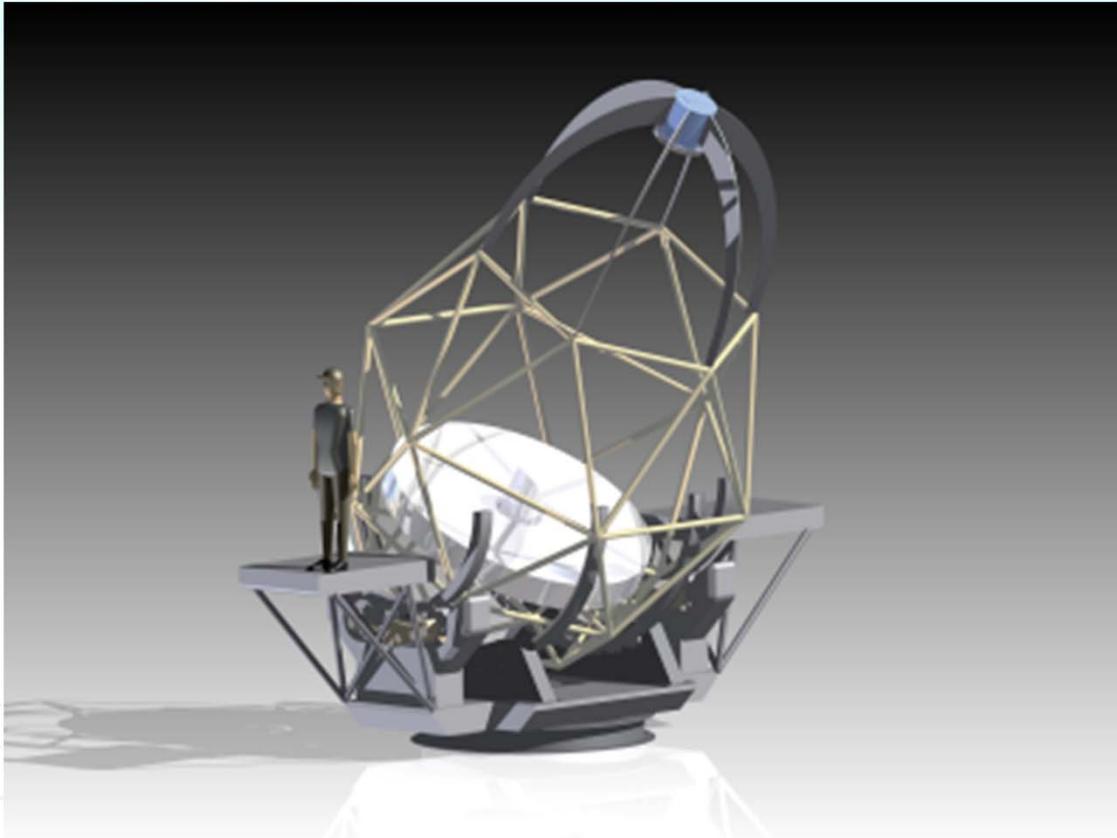


Aurora cameras

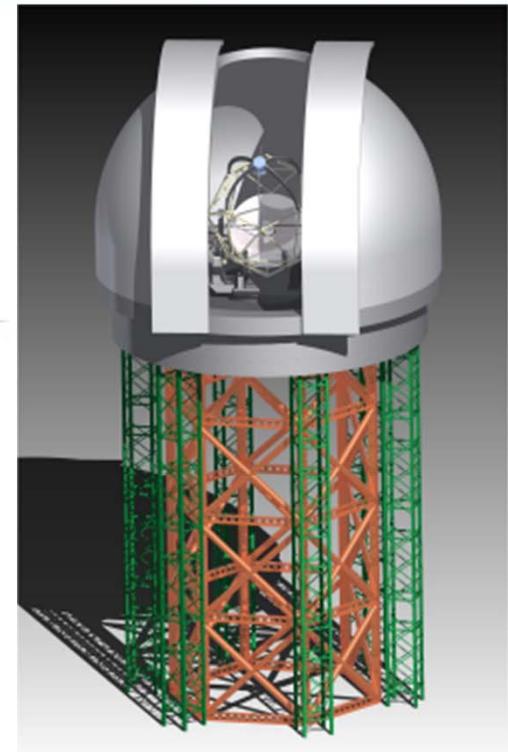


2.5m Infrared Telescope

Ultra-lightweight telescope mount (Kurita+2009)



- Three-Color Infrared Camera (1~5 μm)
+ simple multi-slit spectrograph (option)
- Heterodyne spectrograph (R~100000)
10 μm 、17 μm 、(30 μm)



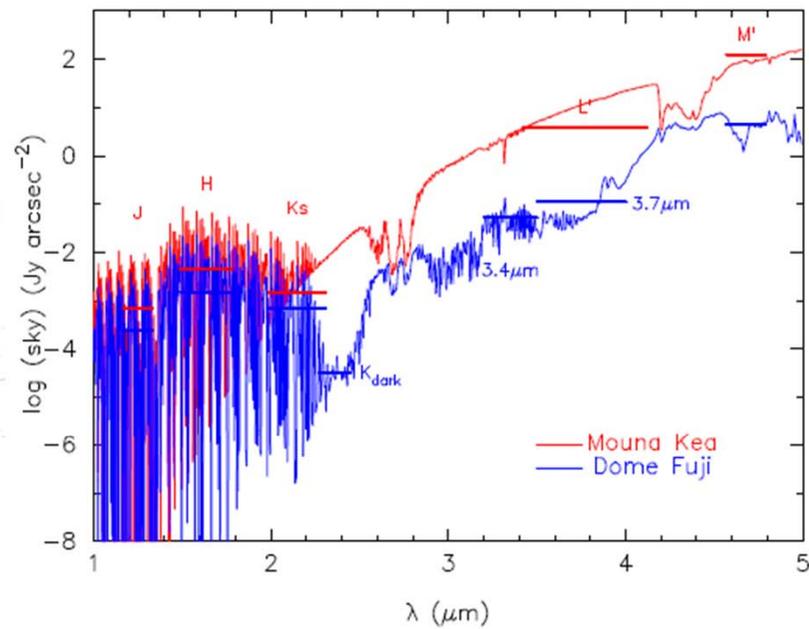
Good Reasons

- Clear sky (>65% photometric), weak wind
- **Cold atmosphere: dark infrared sky (50 – 100 times darker)**
- The free-atmosphere seeing 0.2", the best in the world
- Dry atmosphere: 0.14mmPWV (~10 times lower) in winter
- Stable transparency
- The atmospheric boundary layer is only 11m or lower
- Long periods of uninterrupted darkness for months

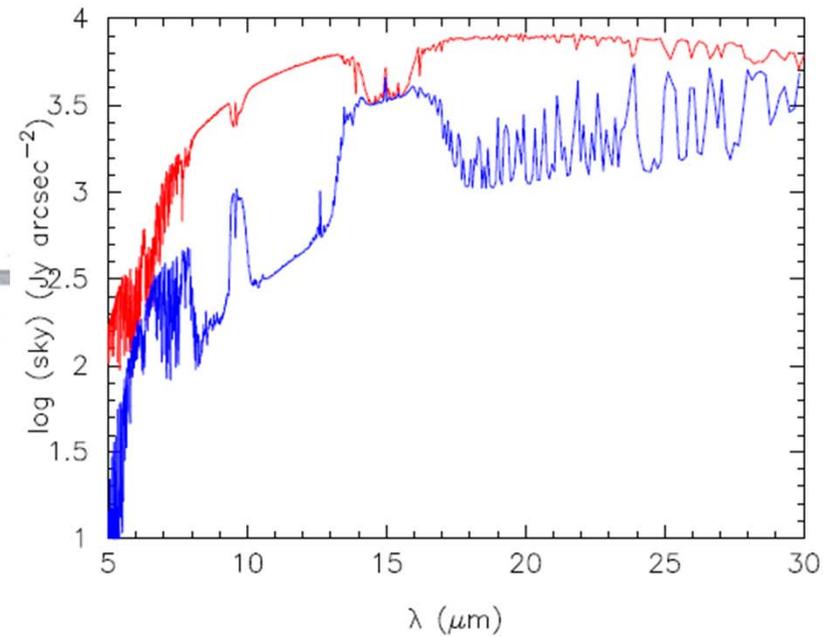
Low sky background

		altitude	temperature
blue	Dome Fuji	3810m	-70°C
red	Mounakea	4200m	0°C

Near-infrared



Mid-infrared



Good Reasons

- Cold atmosphere: dark infrared sky (50 – 100 times darker)
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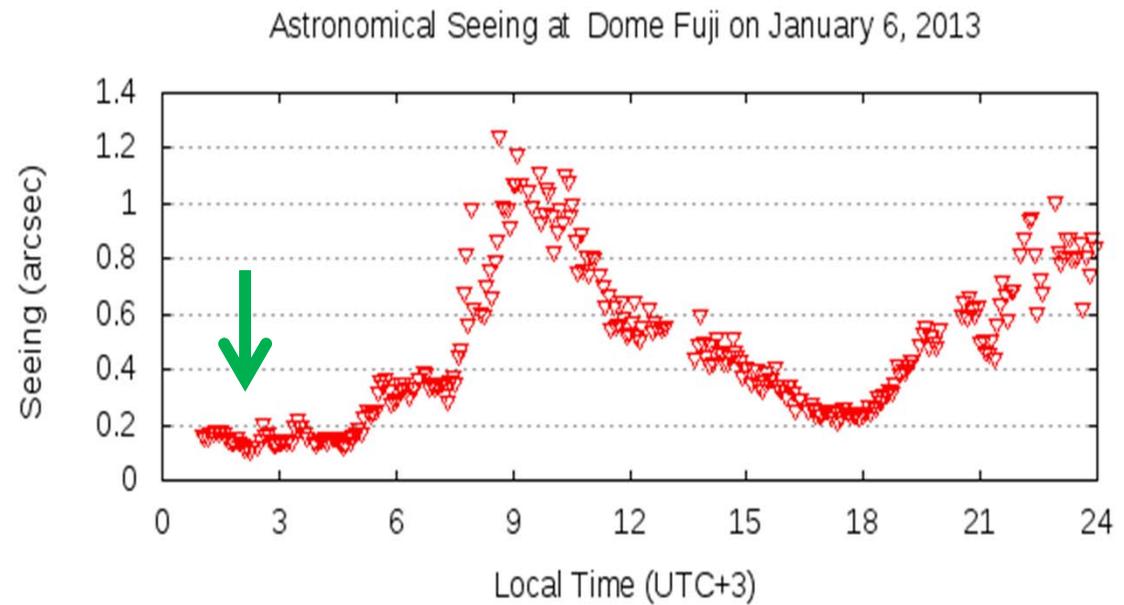
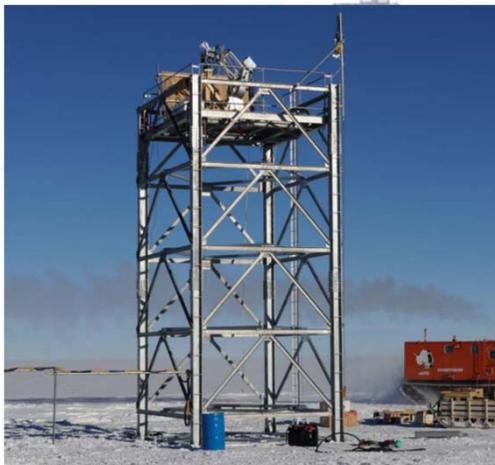
Excellent seeing



Okita's talk



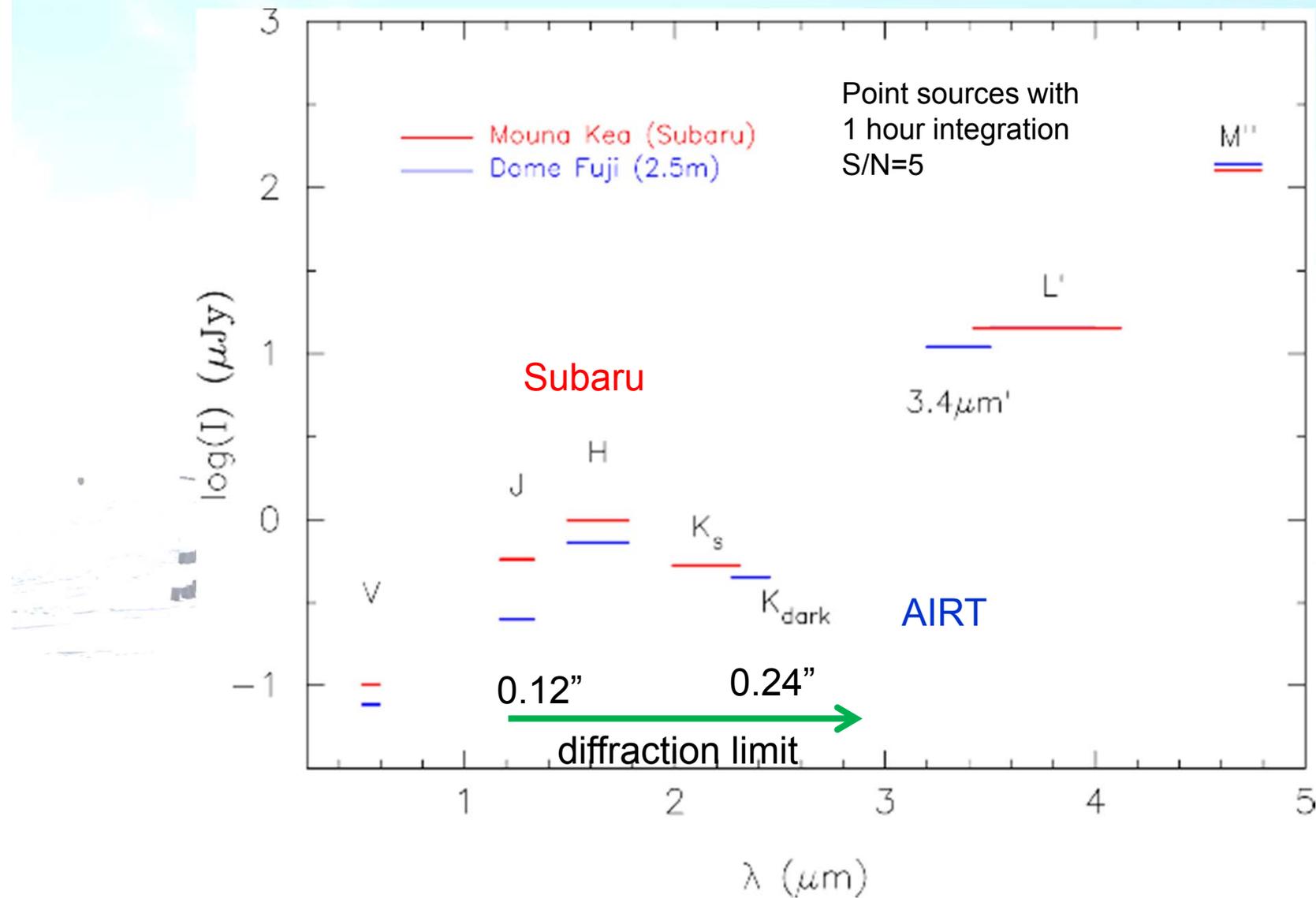
11m



$\sim 0.2''$ ($\lambda = 0.47 \mu\text{m}$)

Okita+ 2013

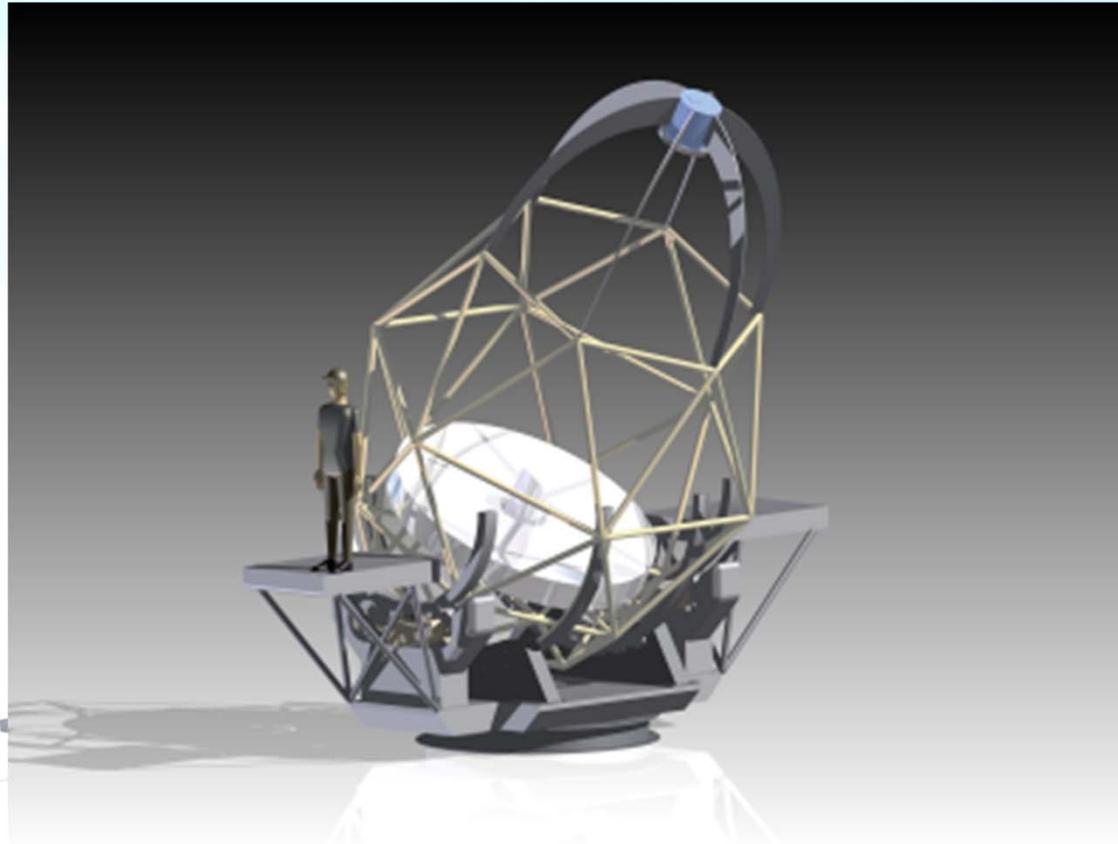
Comparison of Sensitivities - Subaru and AIRT



- ✓ High sensitivity of Subaru telescope
- ✓ High quality of HST

Unique opportunity for deep, high photometric and spatial precision astronomy

In near-infrared,

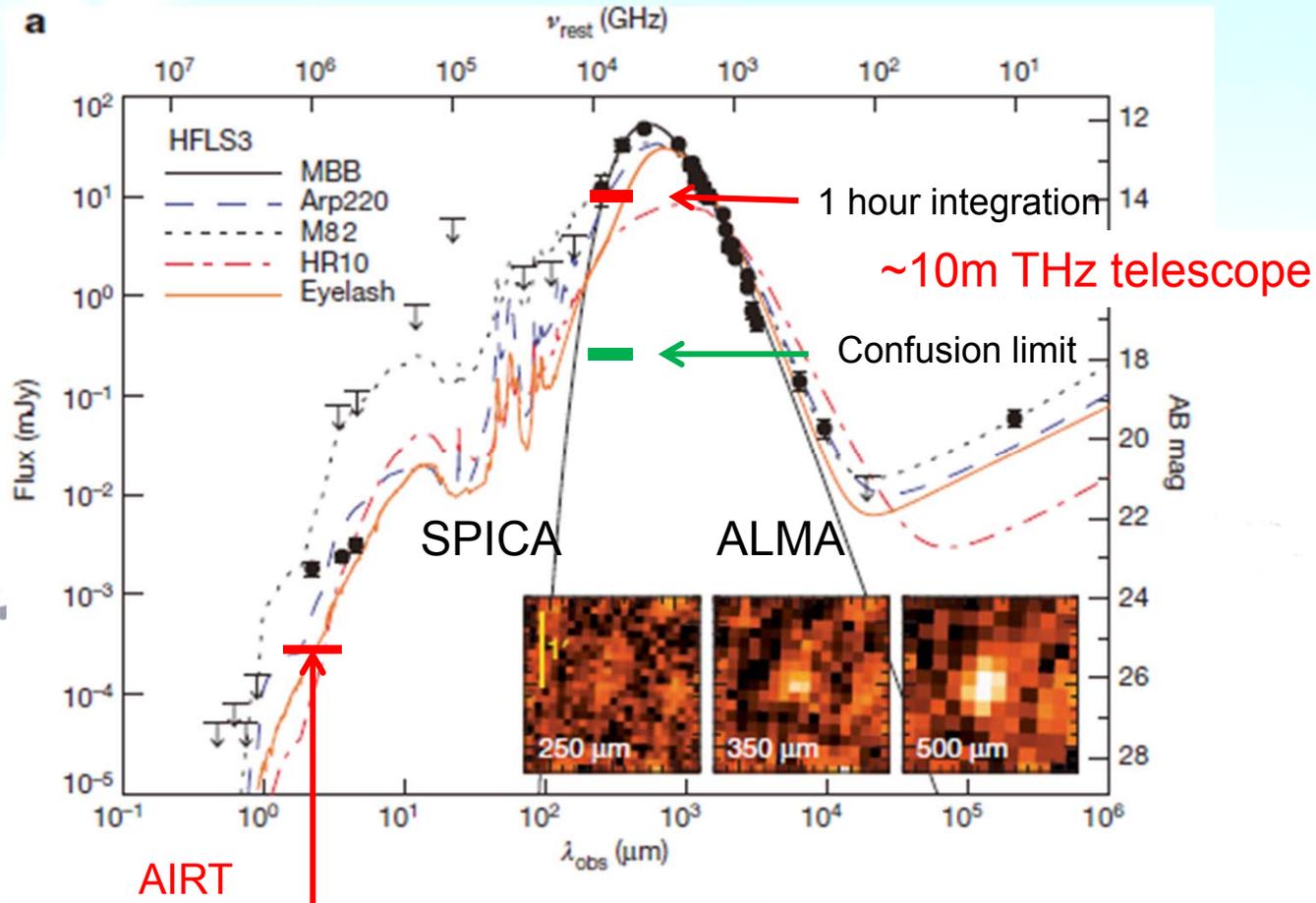


“Big science” with small telescope

➤ Dusty Star Burst Galaxies

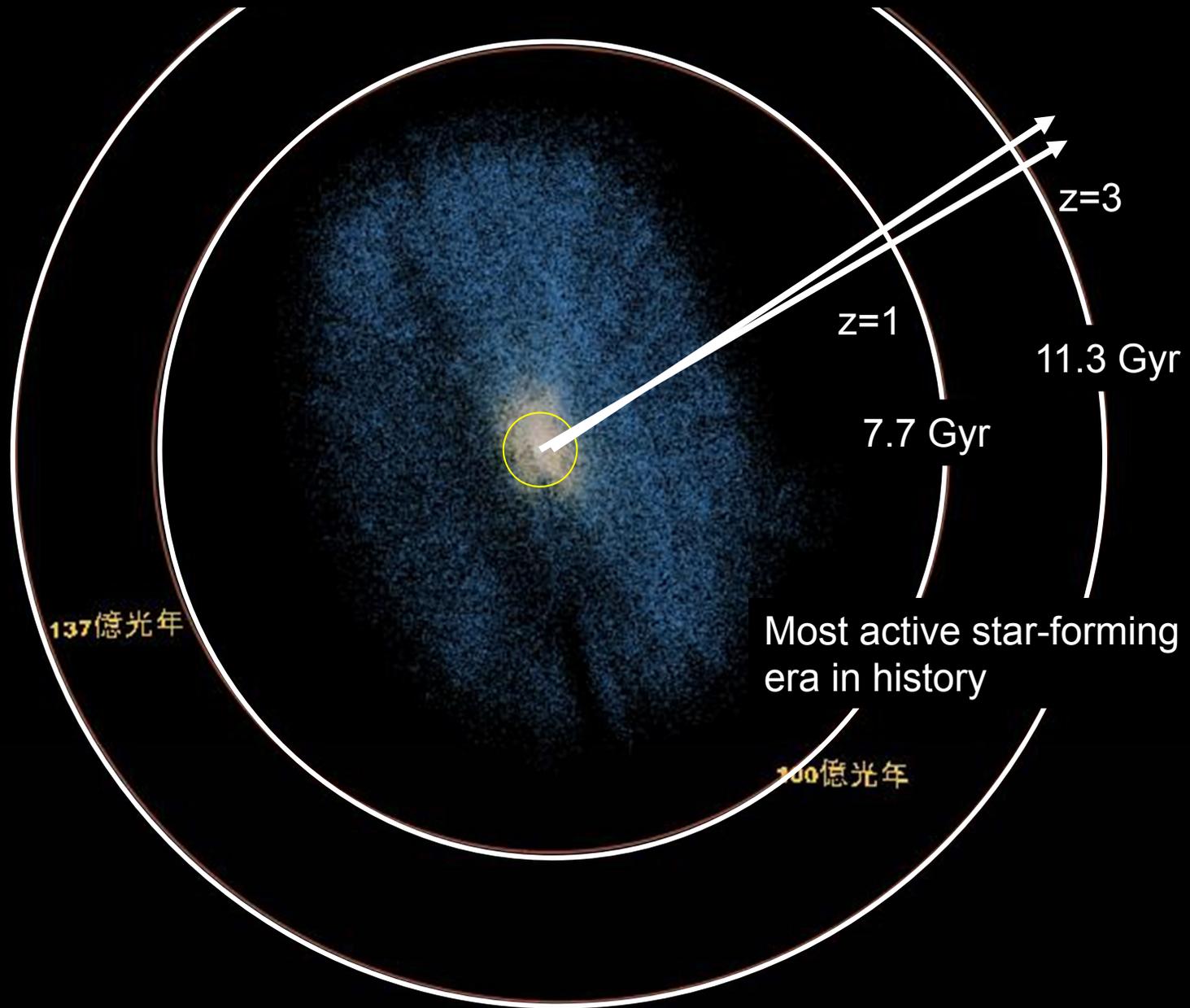
Herschel galaxy at $z=6.5$

Riechers+ 2013



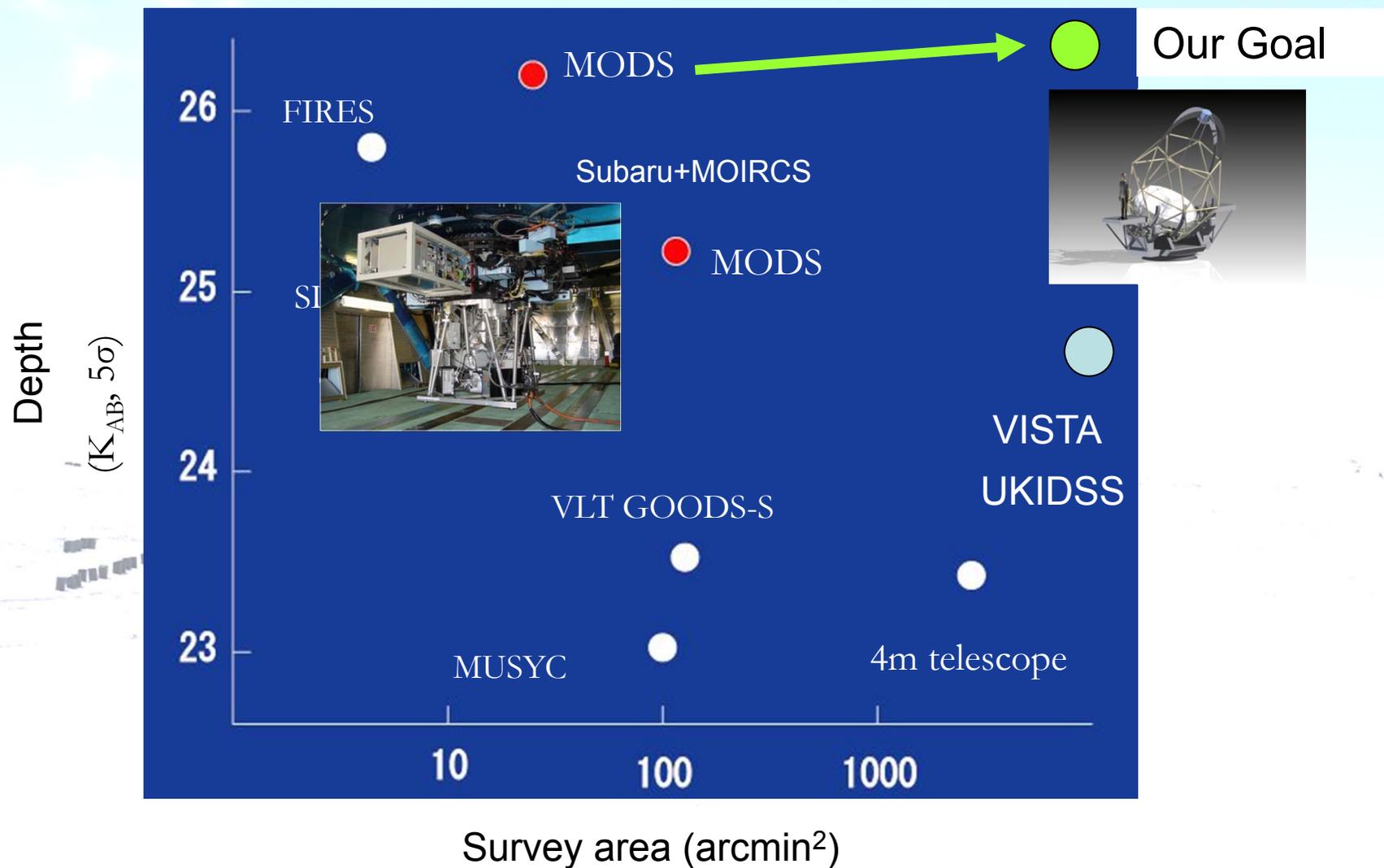
2.4 μm , S/N=5 (1 hour)

Near-infrared wide area survey

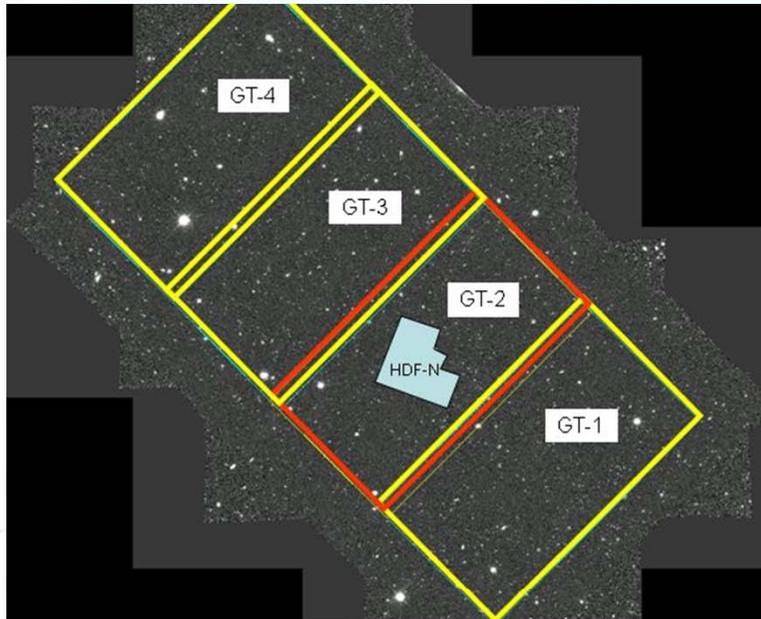


Widest and Deepest High-Redshift Galaxy Survey in K band

Complete samples of $10^9 M_{\text{sun}}$ at $z \sim 3$

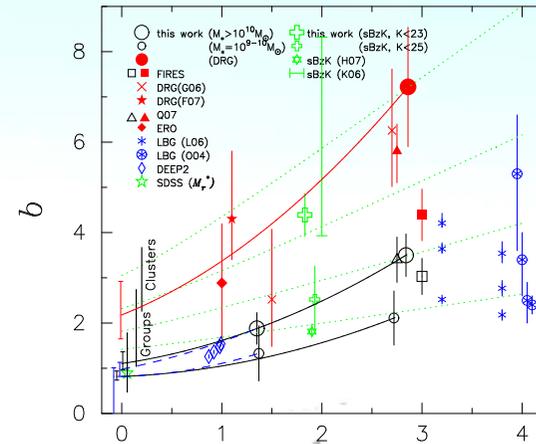


➤ Deepest K-band galaxies with MOIRCS + HST WFC3 data



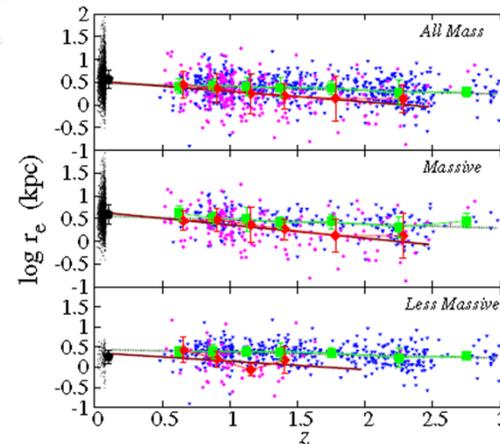
GOODS-N region

Clustering evolution of low mass galaxies



Ichikawa+ 2007

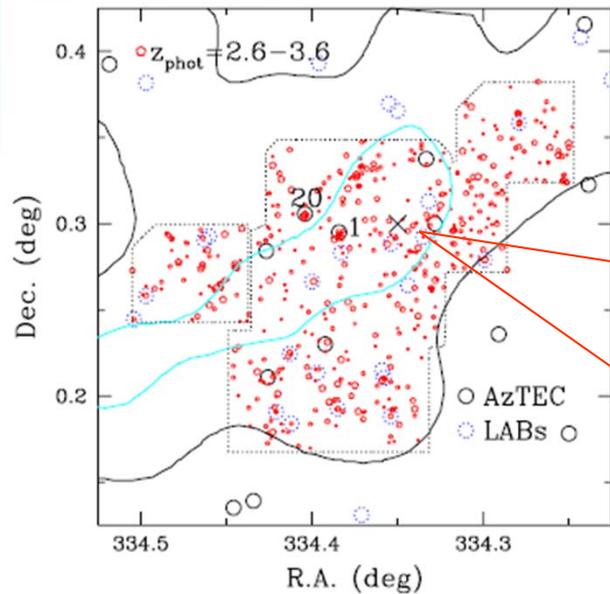
Size evolution



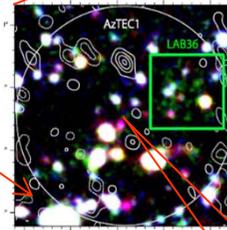
Morishita+ 2013
(submitted)

Large scale structure of galaxy populations

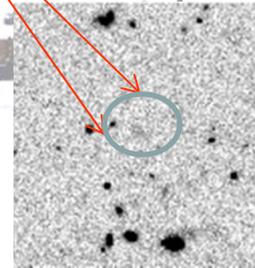
SSA22 ($z \sim 3$) (Uchimoto+2012)



- Massive quiescent and SF galaxies
- Low mass galaxies ($\sim 10^9 M_{\text{sun}}$)
- THz galaxies
- Morphology



100Mpc



Proto-Quasar?

AzTEC1

Tamura+ 2010

THz + near-infrared

Observations will be completed in one season

100Mpc

Good Reasons

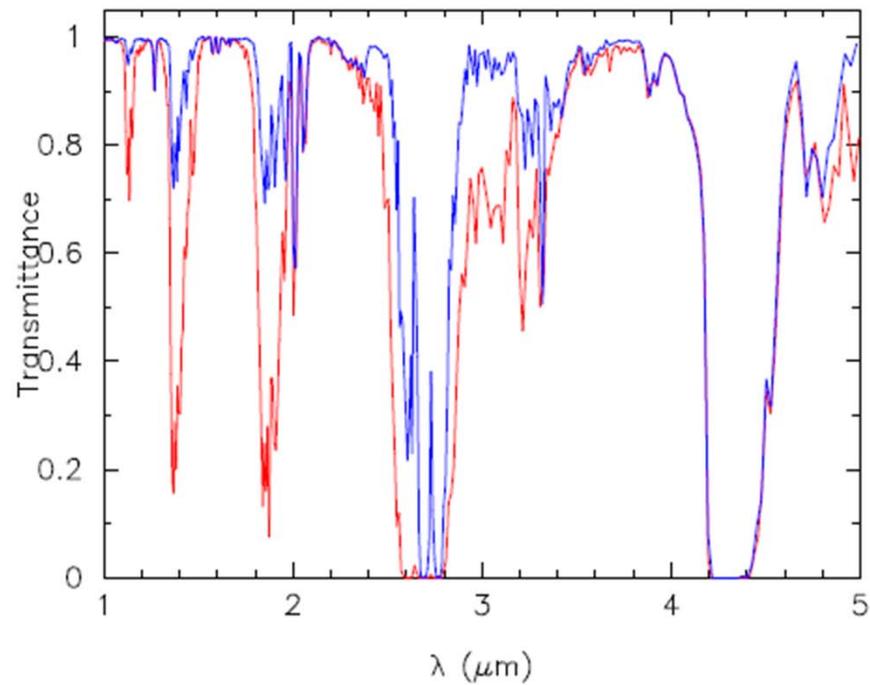
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- Stable transmittance
- Long periods of uninterrupted darkness for months



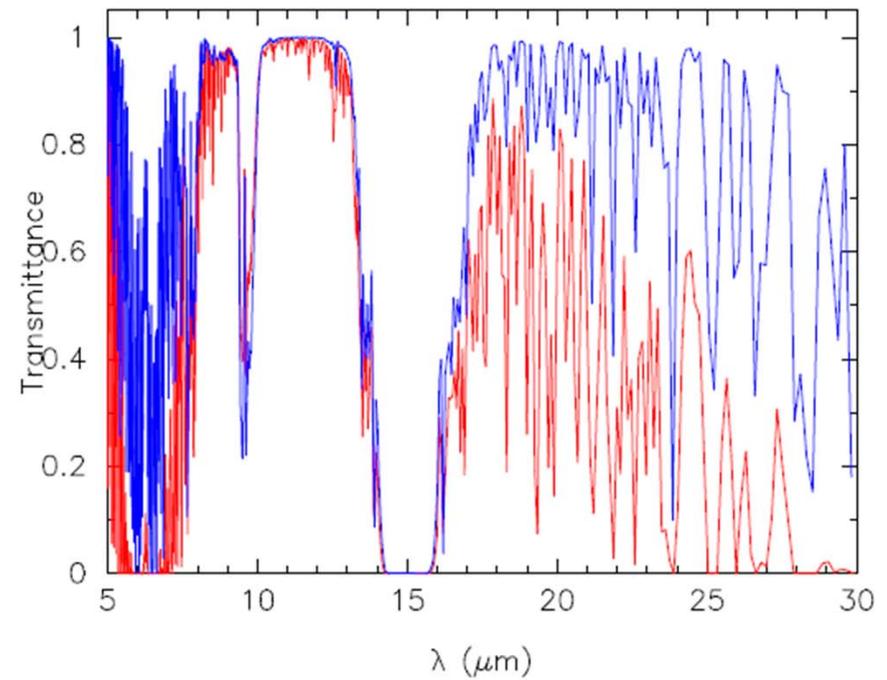
High Transmittance

		altitude	PWV
blue	Dome Fuji	3810m	0.2mm
red	Mounakea	4200m	1mm

Near-infrared



Mid-infrared



Good Reasons

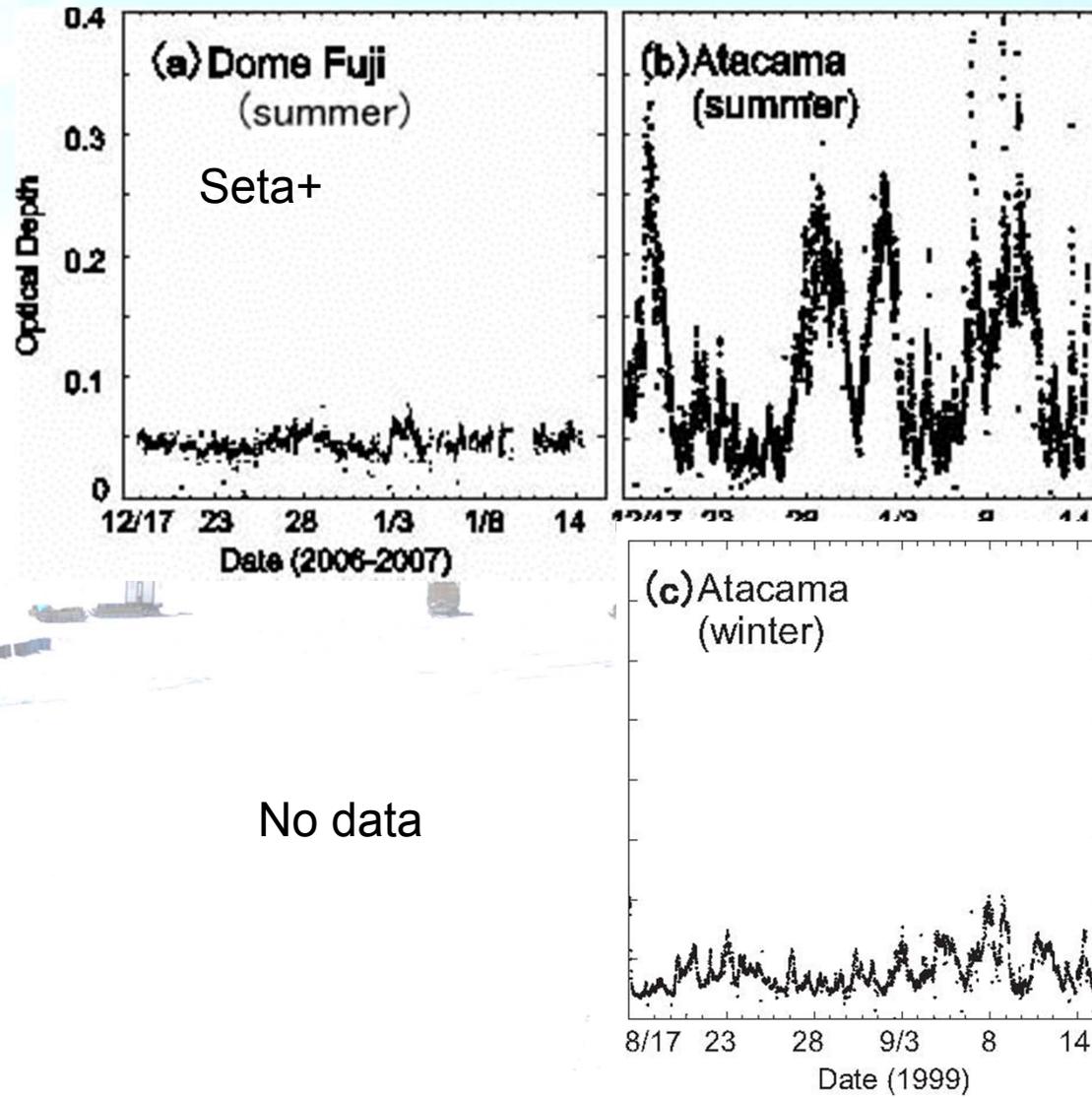
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Stability of Optical Depth in Summer

Dome F

ALMA site



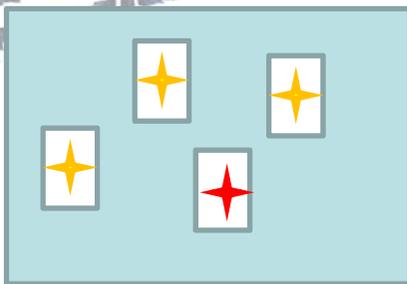
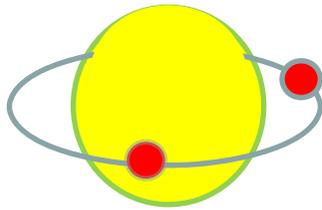
No data

➤ water-dominated atmosphere on super-earths

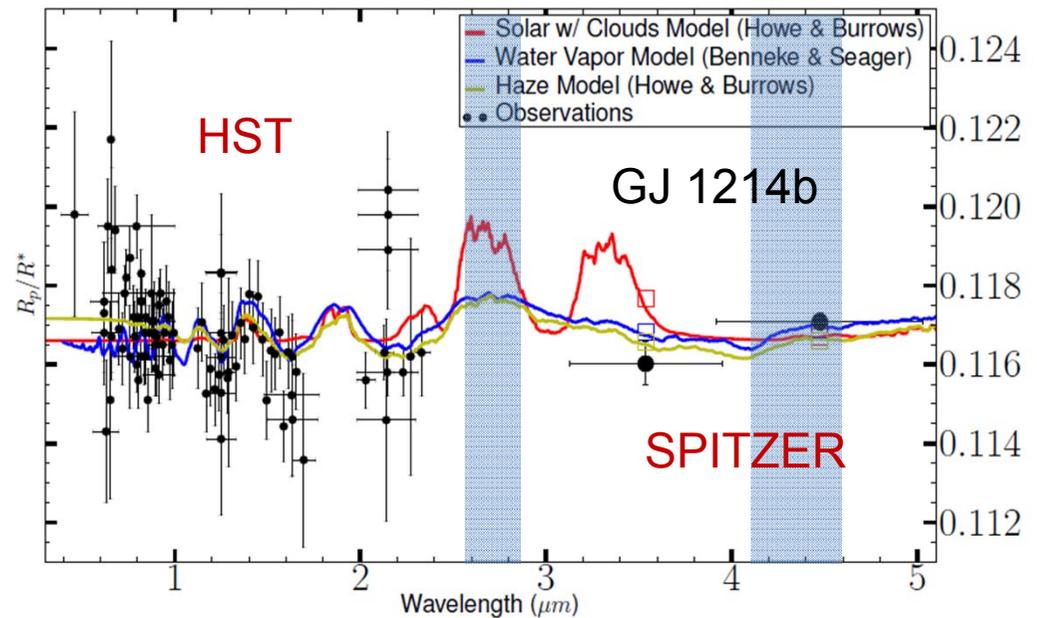
Spectroscopic transit observations with low resolution $\lambda/\Delta\lambda \sim 100$ on multi-slits

Minimum effect of terrestrial water vapor in Antarctica

Transit and secondary eclipse



~10 reference stars in large slits



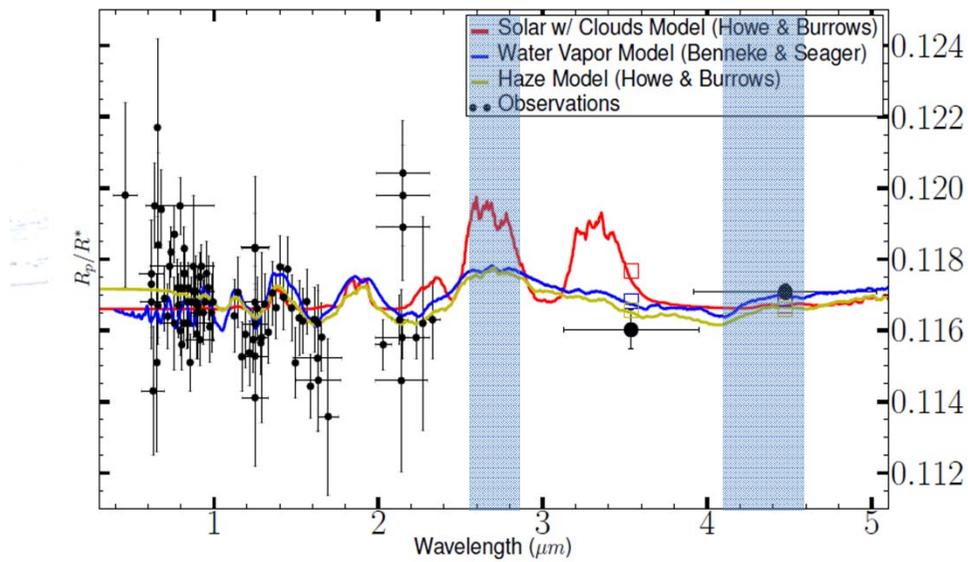
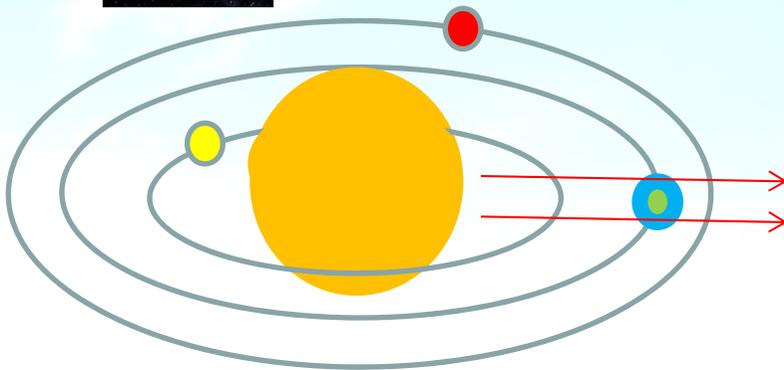
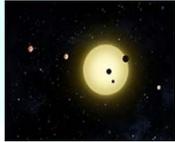
many molecular lines ($\text{H}_2\text{O}, \text{CO}_2, \text{CH}_4 \dots$)

Good Reasons

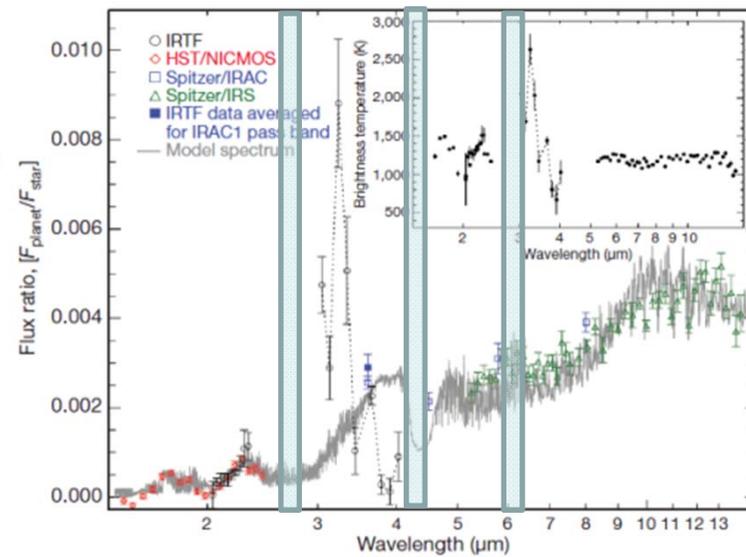
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➤ Continuous observations of multiple systems



Super earth



Hot Jupiter

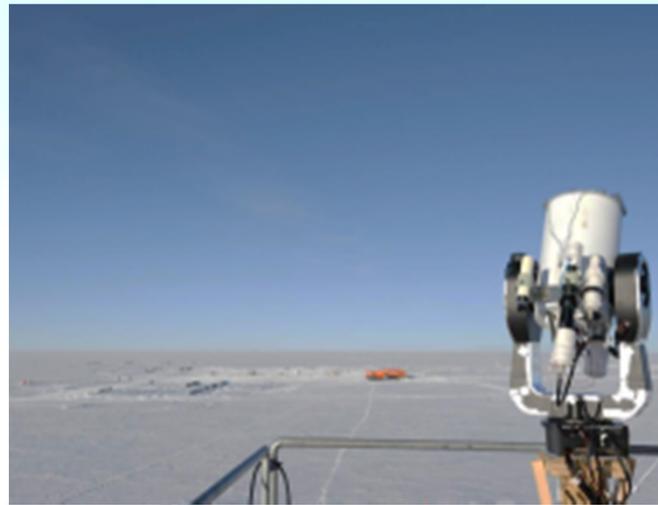
Candidates of multiple systems

TESS (2017)

Bright G, K, M type stars



CCD camera at Dome-F



Collaboration with Dome A

50cm Schmidt x3