# **Event Horizon Telescope Next Steps:** EHT 2020s, ngEHT, and Space Extension

**MIT Haystack Observatory** NRAO Jansky Fellow —> MIT Research Scientist



Event Horizon Telescope

## Kazu Akiyama





## First M87 Results: Where are we now?

- Einstein's GR has passed a new test at an extremely strong gravitational field
- The strongest evidence for the presence of a supermassive black hole
- The M87 central black hole is most likely spinning
- An AGN and associated jet are powered by a supermassive black hole
- The stellar dynamical mass is correct (6.5 billion masses)
- Day-to-day variations on horizon scale

## Dawn of a New Era of Black Hole Astrophysics

## **Testing General Relativity**

M87 Mass: the mass uncertainty is large





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## **Testing General Relativity**

M87 Mass: the mass uncertainty is large





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- Separate the photon ring emission from the surrounding accretion flow and jet

### EHTC+19 Paper I





## **Testing General Relativity**

M87 Mass: the mass uncertainty is large



![](_page_4_Picture_4.jpeg)

- Separate the photon ring emission from the surrounding accretion flow and jet
- Short time window (< a week): may significantly be affected by transient feature

![](_page_4_Picture_11.jpeg)

## **BH** Magnetosphere, Accretion and Jet

## Tracking evolving features in M87

![](_page_5_Figure_3.jpeg)

![](_page_5_Picture_4.jpeg)

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### EHTC+19 Paper I

![](_page_5_Picture_9.jpeg)

## **BH Magnetosphere, Accretion and Jet**

- Tracking evolving features in M87
- Black Hole Magnetosphere: Magnetic Flux (SANE / MAD)

![](_page_6_Picture_4.jpeg)

![](_page_6_Picture_5.jpeg)

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EHTC+19 Paper I

## **BH Magnetosphere, Accretion and Jet**

Tracking evolving features in M87 Black Hole Magnetosphere: Magnetic Flux (SANE / MAD) Jet launching: how is the energy extracted from the black hole?

![](_page_7_Picture_3.jpeg)

![](_page_7_Picture_4.jpeg)

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# M87 Polarimetry: Coming Soon

![](_page_8_Figure_1.jpeg)

![](_page_8_Picture_2.jpeg)

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## Akiyama et al. 2017, Chael et al. 2016

![](_page_8_Figure_6.jpeg)

# Sgr A\*: Another Horizon-scale Target

## **Best target for the GR test**

Mass & Distance are accurately measured \_ (GRAVITY Collaboration+2018)

## **Current challenges**

- Time variations: on minutes time scales
- Interstellar Scattering: much less dominant

![](_page_9_Picture_6.jpeg)

![](_page_9_Picture_7.jpeg)

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![](_page_9_Figure_10.jpeg)

### GRAVITY Collaboration+2018

## Weather forecast: Sgr A\*

## An illustrative full-closure imaging simulation with EHT 2017 array

![](_page_10_Picture_2.jpeg)

![](_page_10_Picture_3.jpeg)

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![](_page_10_Picture_6.jpeg)

### Model: Broderick & Loeb 2006, Imaging: Kotaro Moriyama

![](_page_10_Picture_8.jpeg)

# EHT2020s: EHT within 3 years from now

![](_page_11_Picture_1.jpeg)

Doeleman et al. 2019, Astro2020 white paper

2017: 230 GHz, 32 Gbps, 8 stations @ 6 sites 2023: 230 & 345 GHz, 64 Gbps, 12 stations @ 10 sites

![](_page_11_Picture_4.jpeg)

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## 2018:

- **Bandwidth:** 32 Gbps -> 64 Gbps -
- **New Station:** \_
  - GLT (under commissioning)

## 2020:

**New Stations:** GLT (official) NOEMA & Kitt Peak (KP)

## 2023:

- **New Station:** Owens Valley (OVRO)
- 345 GHz Capability

![](_page_11_Figure_16.jpeg)

![](_page_11_Figure_17.jpeg)

![](_page_11_Figure_18.jpeg)

![](_page_11_Figure_19.jpeg)

## EHT 2020s: Deeper, Shaper & Multi-frequency Images

![](_page_12_Picture_1.jpeg)

### 230 GHz EHT2017+GLT

Further new capabilities: Faraday Rotation Imaging

Active developments of multi-frequency (Chael+) and multi-scale (Akiyama+) Imaging

![](_page_12_Picture_5.jpeg)

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![](_page_12_Picture_8.jpeg)

### 230+345 GHz EHT2017+GLT+KP+NOEMA+OVRO

Simulations: Andrew Chael, Imaging: Kazu Akiyama

![](_page_12_Picture_12.jpeg)

![](_page_12_Picture_13.jpeg)

![](_page_12_Picture_14.jpeg)

![](_page_12_Picture_15.jpeg)

# EHT 2020s: Tracking Evolving Features

![](_page_13_Picture_1.jpeg)

### Regular monitoring observation capabilities on weekly scales

![](_page_13_Picture_3.jpeg)

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Kazu Akiyama, "Active Galactic Nucleus Jets in the Event Horizon Telescope Era", Tohoku University, 2020/01/20 (Mon)

![](_page_13_Figure_6.jpeg)

Simulations: Charles Gammie, George Wong et al., Imaging: Michael Johnson

![](_page_13_Figure_8.jpeg)

# EHT 2020s: Precision Black Hole Astrophysics

![](_page_14_Picture_1.jpeg)

## Mass? Spin? Accretion flow types? Viewing Geometry?

![](_page_14_Picture_3.jpeg)

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https://www.youtube.com/watch?v=0ymmnHlnDVY

![](_page_14_Picture_8.jpeg)

![](_page_14_Figure_9.jpeg)

![](_page_14_Picture_10.jpeg)

# EHT 2020s: Precision Black Hole Astrophysics

## Are black hole images confusing for scientists and/or AI?

![](_page_15_Picture_2.jpeg)

![](_page_15_Figure_3.jpeg)

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https://barkpost.com/humor/doodle-or-fried-chicken-twitter/

![](_page_15_Picture_9.jpeg)

![](_page_15_Picture_10.jpeg)

![](_page_15_Picture_11.jpeg)

# EHT 2020s: Precision Black Hole Astrophysics

Van der Gucht et al. 2019

![](_page_16_Figure_2.jpeg)

![](_page_16_Picture_3.jpeg)

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Current forecast: Horizon-scale images are much less confusing!

![](_page_16_Figure_7.jpeg)

# next generation Event Horizon Telescope (2023 -)

![](_page_17_Picture_1.jpeg)

## **Phase I:** 2019-2023 (Array Design Phase + MIT Haystack as a new site) **Phase II:** 2023- (Constructions of several new telescopes)

![](_page_17_Picture_3.jpeg)

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![](_page_17_Picture_6.jpeg)

# M87 ngEHT images: ~10 years from now

![](_page_18_Picture_1.jpeg)

Blackburn et al. 2019; Doeleman et al. 2019 (Astro2020 Decadal Survey White Papers)

![](_page_18_Picture_3.jpeg)

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![](_page_18_Picture_6.jpeg)

# ngEHT: far more details

![](_page_19_Picture_1.jpeg)

### Astro2020 APC White Paper Studying Black Holes on Horizon Scales with **VLBI** Ground Arrays

Lindy Blackburn<sup>1,2,\*</sup> Sheperd Doeleman<sup>1,2,\*</sup>, Jason Dexter<sup>12</sup>, José L. Gómez<sup>16</sup>, Michael D. Johnson<sup>1,2</sup>, Daniel C. Palumbo<sup>1,2</sup>, Jonathan Weintroub<sup>1,2</sup>, Joseph R. Farah<sup>1,2,21</sup>, Vincent Fish<sup>4</sup>, Laurent Loinard<sup>18,19</sup>, Colin Lonsdale<sup>4</sup>, Gopal Narayanan<sup>28</sup>, Nimesh A. Patel<sup>2</sup>, Dominic W. Pesce<sup>1,2</sup>, Alexander Raymond<sup>1,2</sup>, Remo Tilanus<sup>17,22,23</sup>, Maciek Wielgus<sup>1,2</sup>, Kazunori Akiyama<sup>1,3,4,5</sup>, Geoffrey Bower<sup>6</sup>, Avery Broderick<sup>7,8,9</sup>, Roger Deane<sup>10,11</sup>, Christian Michael Fromm<sup>13</sup>, Charles Gammie<sup>14,15</sup>, Roman Gold<sup>13</sup>, Michael Janssen<sup>17</sup>, Tomohisa Kawashima<sup>4</sup>, Thomas Krichbaum<sup>29</sup>, Daniel P. Marrone<sup>20</sup>, Lynn D. Matthews<sup>4</sup>, Yosuke Mizuno<sup>13</sup>, Luciano Rezzolla<sup>13</sup>, Freek Roelofs<sup>17</sup>, Eduardo Ros<sup>29</sup>, Tuomas K. Savolainen<sup>29,30,31</sup>, Feng Yuan<sup>24,25,26</sup>, Guangyao Zhao<sup>27</sup>

![](_page_19_Picture_4.jpeg)

<sup>1</sup> Black Hole Initiative at Harvard University, 20 Garden Street, Cam- and Particle Physics (IMAPP), Radboud University, P.O. Box 9010, bridge, MA 02138, USA 6500 GL Nijmegen, The Netherlands Event Horizon Telescope

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### Astro2020 APC WP

Blackburn+

# EHT beyond the Earth: Why do we need?

## **Ground limitations:**

## **Telescope Coverages**

Limiting snapshot imaging of Sgr A\* and multi-epoch observations

## **Angular resolution**

- The maximum diameter = the Earth diameter

![](_page_20_Picture_10.jpeg)

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## Not so many sites capable of sub-mm VLBI observations Snapshot coverages will be highly limited by Earth rotation

Frequency / wavelength: limited by atmosphere (up to ~350 GHz)

Sub-Rs scale imaging of M87/Sgr A\*, more horizon-scale targets

## EHT beyond the Earth: Designing towards late 2020

![](_page_21_Picture_1.jpeg)

![](_page_21_Picture_2.jpeg)

**Event Horizon Telescope** 

Kazu Akiyama, "Active Galactic Nucleus Jets in the Event Horizon Telescope Era", Tohoku University, 2020/01/20 (Mon)

### 1st KISS Workshop, Sep 2019

![](_page_21_Picture_6.jpeg)

![](_page_21_Picture_7.jpeg)

## EHT beyond the Earth: Designing towards late 2020

- Astro 2020 Science and APC white papers (Doeleman+, Haworth+, Pesce+)
- Active International Discussions Sep 2018: 1st International Space VLBI conference @ Netherland Sep 2019: 1st KISS Workshop @ USA (invitation only) Jan 2020: 2nd International Space VLBI conference @ USA (Next week!) June 2020: 2nd KISS Workshop @ USA (invitation only)

![](_page_22_Picture_3.jpeg)

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Kazu Akiyama, "Active Galactic Nucleus Jets in the Event Horizon Telescope Era", Tohoku University, 2020/01/20 (Mon)

![](_page_22_Picture_9.jpeg)

Astro2020 APC White Paper

### **Extremely Long-Baseline Interferometry** with the Origins Space Telescope

DOMINIC W. PESCE<sup>1,2</sup>, KARI HAWORTH<sup>1</sup>, GARY J. MELNICK<sup>1</sup>, LINDY BLACKBURN<sup>1,2</sup>, MACIEK WIELGUS<sup>1,2</sup>, ALEXANDER RAYMOND<sup>1,2</sup>, JONATHAN WEINTROUB<sup>1</sup>, DANIEL C. M. PALUMBO<sup>1,2</sup>, MICHAEL D. JOHNSON<sup>1,2</sup>, SHEPERD S. DOELEMAN<sup>1,2</sup>, DAVID J. JAMES<sup>1,2</sup>

1st KISS Workshop, Sep 2019

![](_page_22_Picture_17.jpeg)

![](_page_22_Picture_18.jpeg)

![](_page_22_Picture_19.jpeg)

# EHT+LEO: Better Snapshot Imaging

![](_page_23_Figure_1.jpeg)

## EHT+MEO/GEO: (sub-)R<sub>s</sub>-scale Imaging of M87 / Sgr A\*

![](_page_24_Figure_1.jpeg)

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### (Kawashima, Kino & Akiyama 2019, ApJ)

![](_page_24_Picture_6.jpeg)

![](_page_24_Picture_7.jpeg)

## EHT+MEO/GEO: Horizon-scale Imaging of other SMBHs

![](_page_25_Figure_1.jpeg)

![](_page_25_Picture_2.jpeg)

Kazu Akiyama, "Active Galactic Nucleus Jets in the Event Horizon Telescope Era", Tohoku University, 2020/01/20 (Mon)

## M104 (Sombrero Galaxy)

![](_page_25_Figure_5.jpeg)

### Fish, Shea & Akiyama 2019

![](_page_25_Figure_7.jpeg)

## Far more accurate shadow measurements?

![](_page_26_Figure_1.jpeg)

![](_page_26_Figure_2.jpeg)

Johnson et al. 2019

![](_page_26_Figure_6.jpeg)

## Far more accurate shadow measurements?

![](_page_27_Figure_1.jpeg)

![](_page_27_Figure_6.jpeg)

# Weather Forecast of the EHT era: Awesome!

## EHT in the next three years:

- Sgr A\* & M87 Polarimetry
- Doubled bandwidth (2018-): Higher sensitivity & Rotation Measure Imaging
- Four new stations until 2023: Higher dynamic range (few/several 100s)
- Higher angular resolution: 345 GHz (0.87 mm)
- Time domain: Monitoring Capability

## mid-late 2020s: Expanding the Array to the ngEHT

Higher Dynamic Range (~1000s), Better sensitivity

## late 2020s-2030s: Expanding the ngEHT to space:

![](_page_28_Picture_11.jpeg)

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- Better snapshot, sub-Rs-scale imaging, more horizon-scale targets

![](_page_28_Picture_18.jpeg)

![](_page_28_Picture_19.jpeg)

![](_page_28_Picture_20.jpeg)